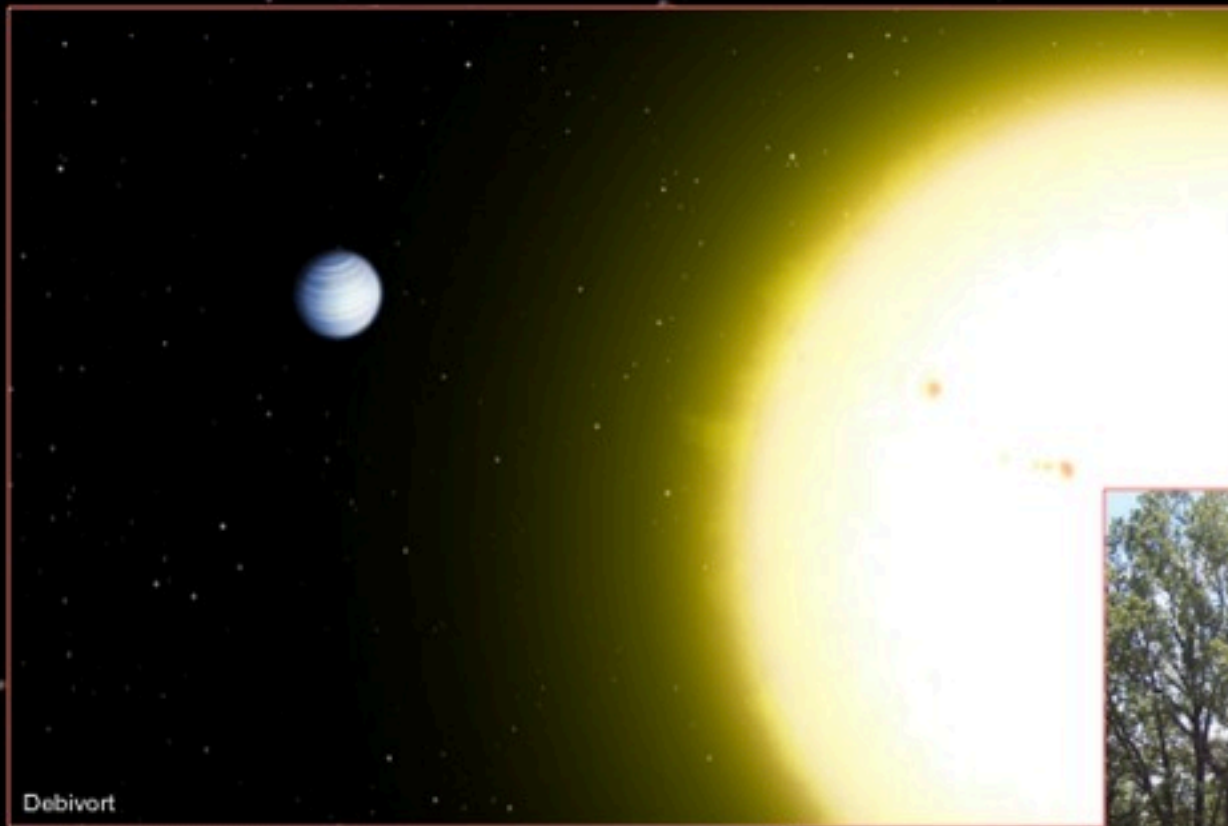


Exoplanet around a Sun-like Star



Debivort

51 Peg b
OHP
France



Eric Ford

1543

1609

1781

1846

1993

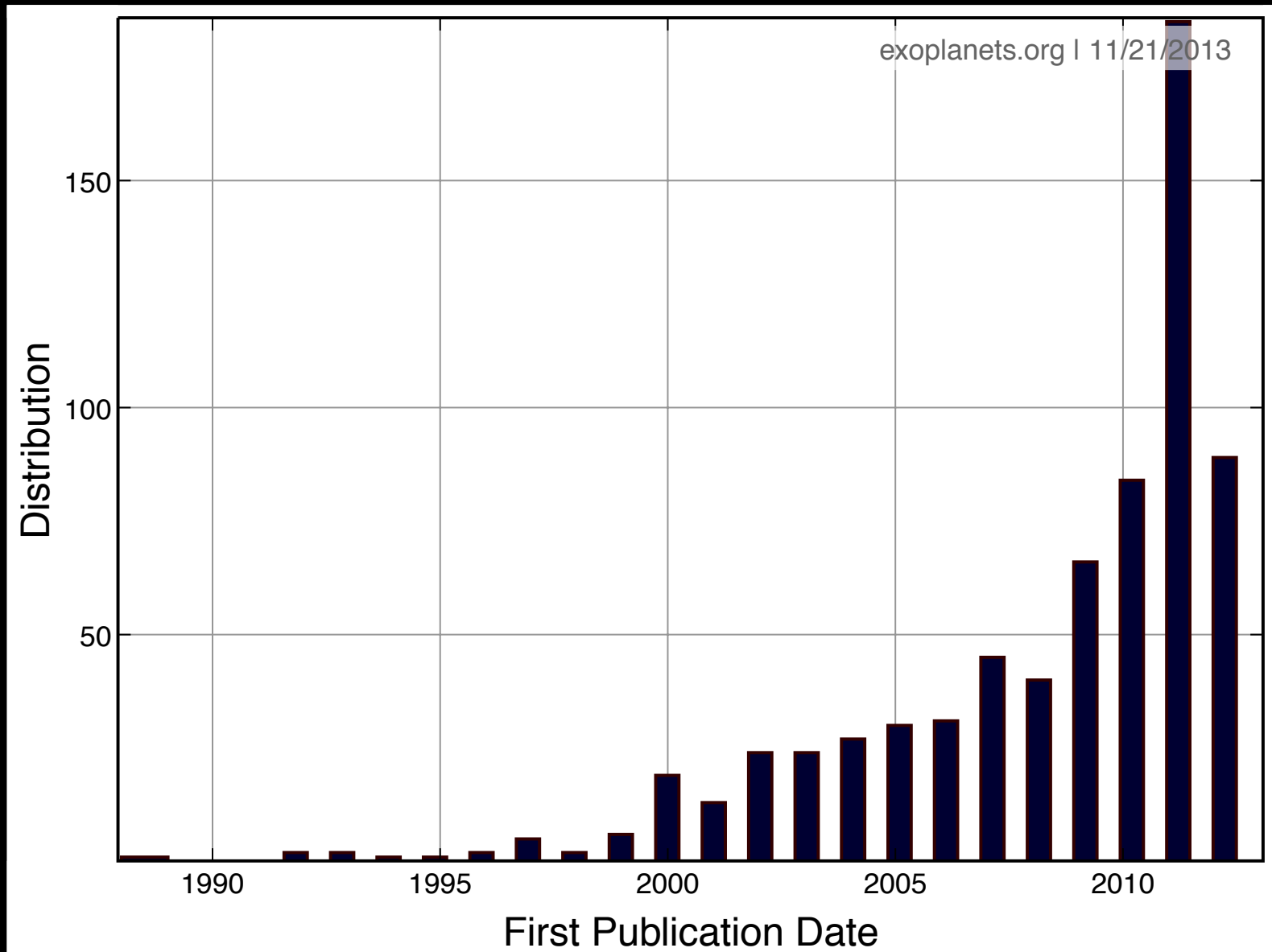
1995

1999

2001

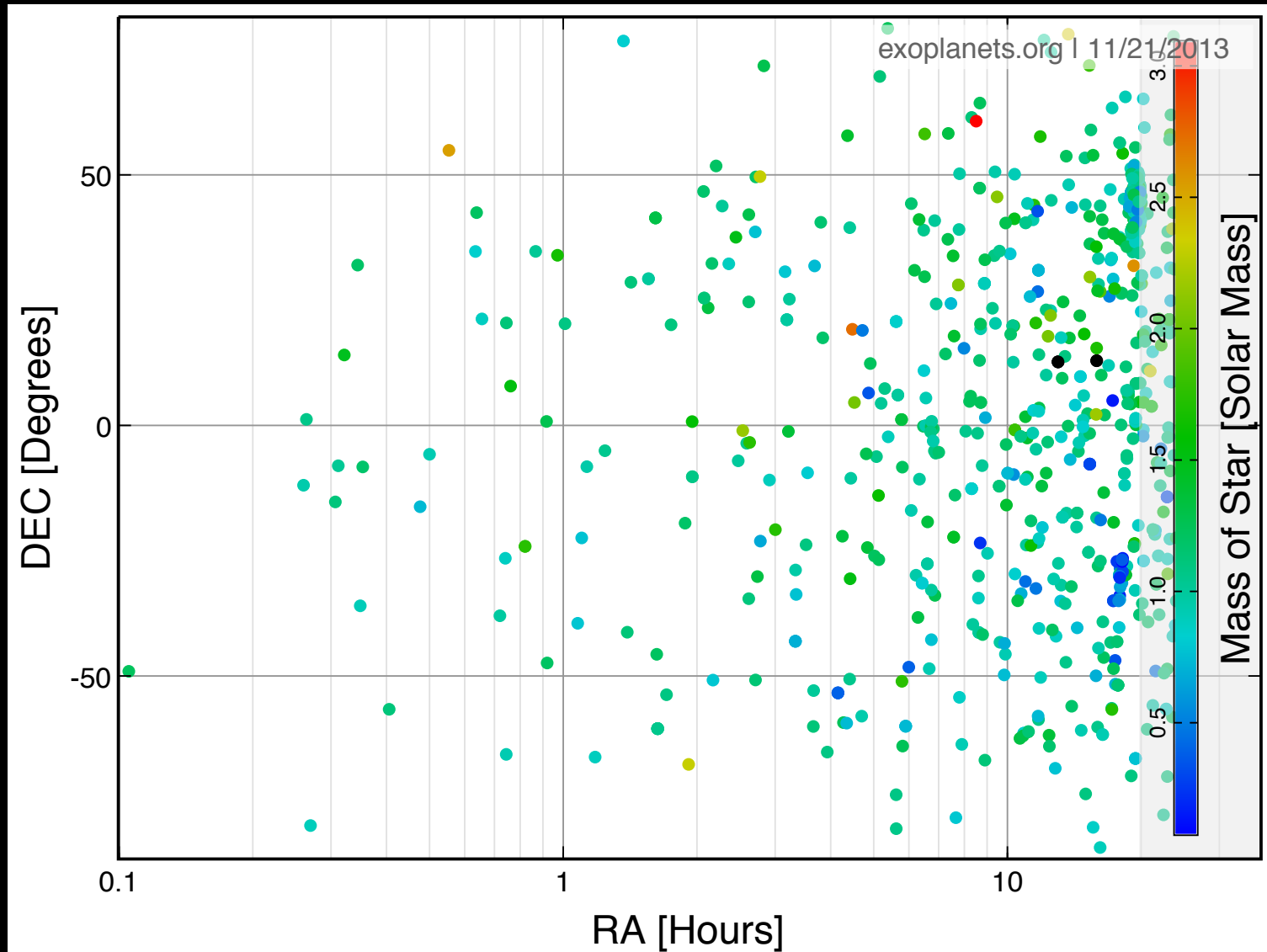
2009

Planet Detection Timeline

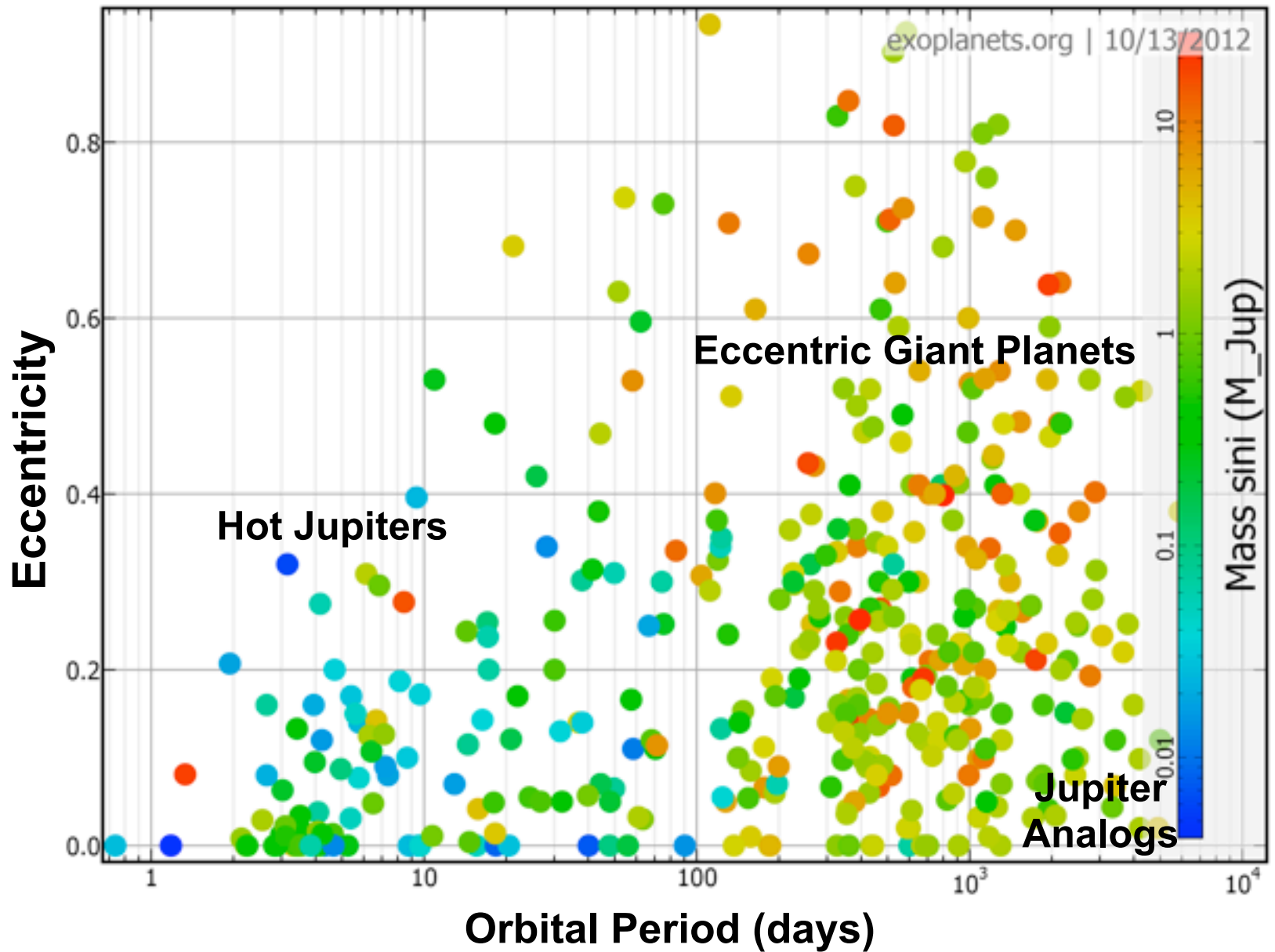


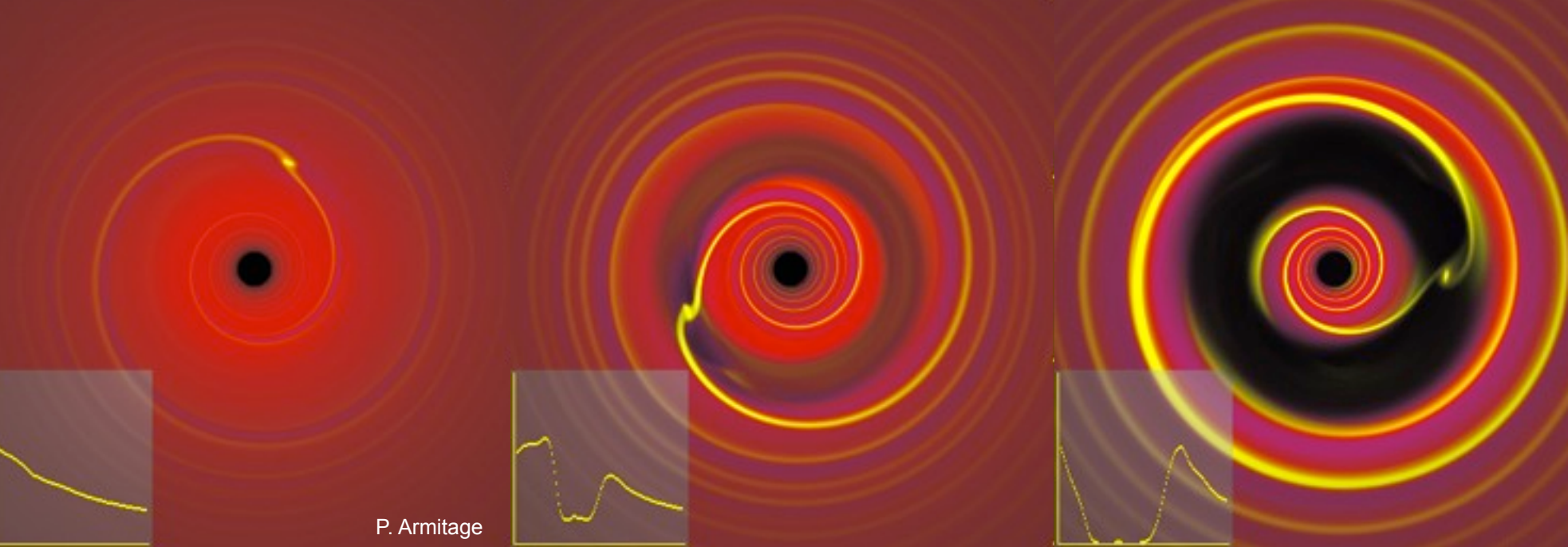
Exoplanet around a Sun-like Star

They are everywhere!

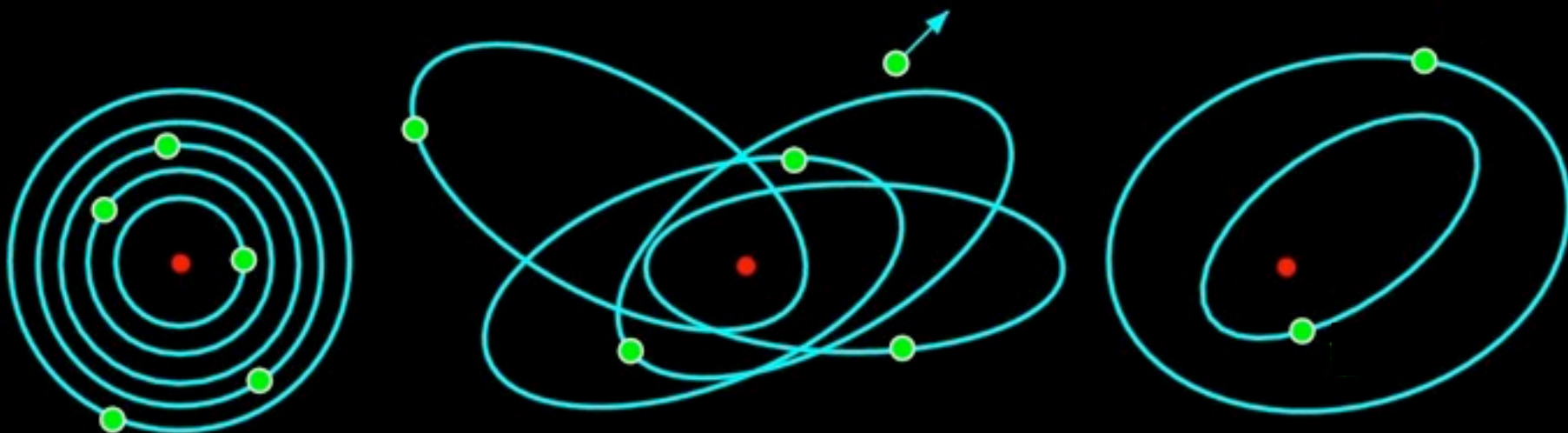


Diversity of Extrasolar Planets





New Theories of Planet Formation



Hot Jupiters via Disk Migration

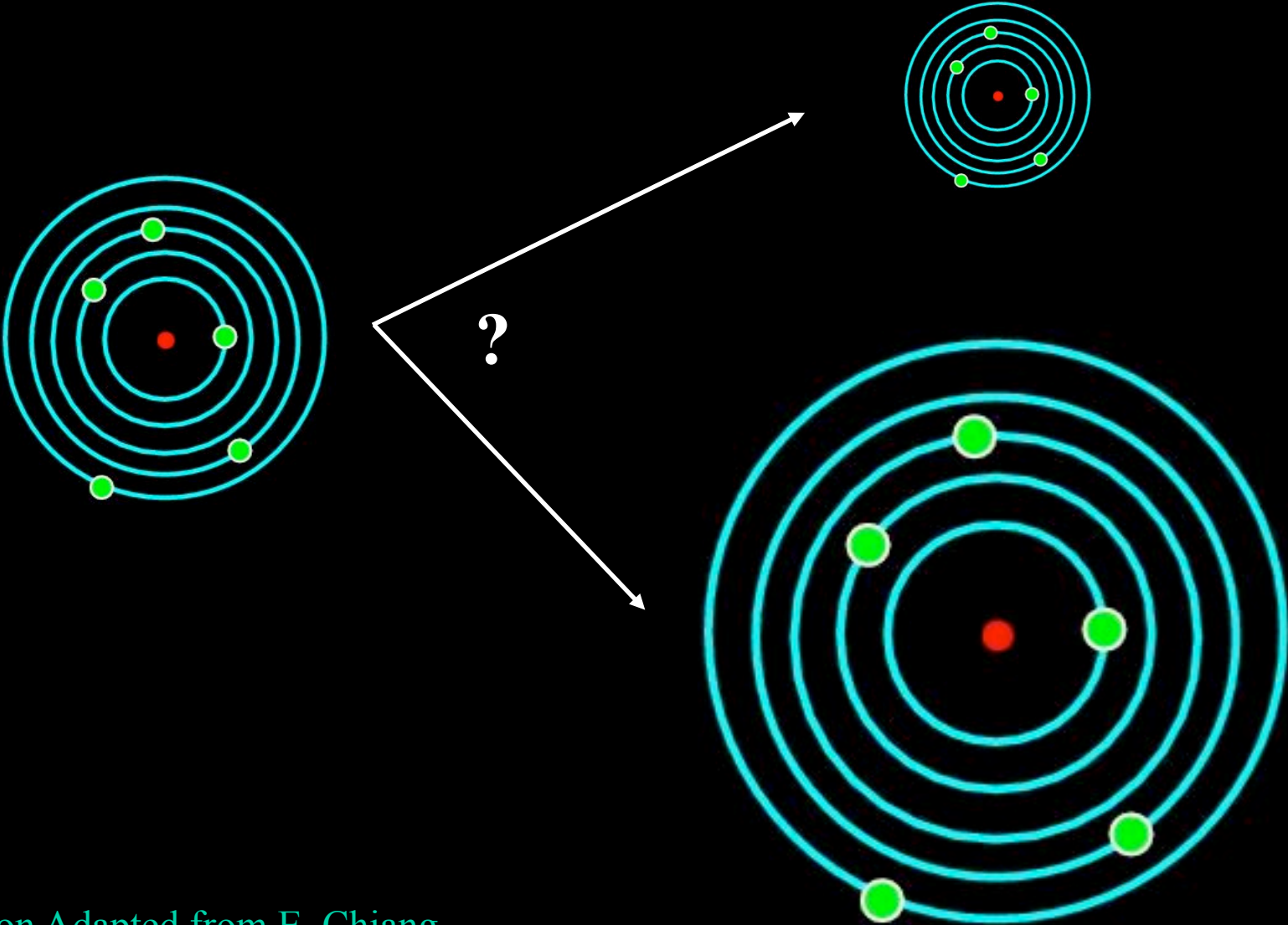
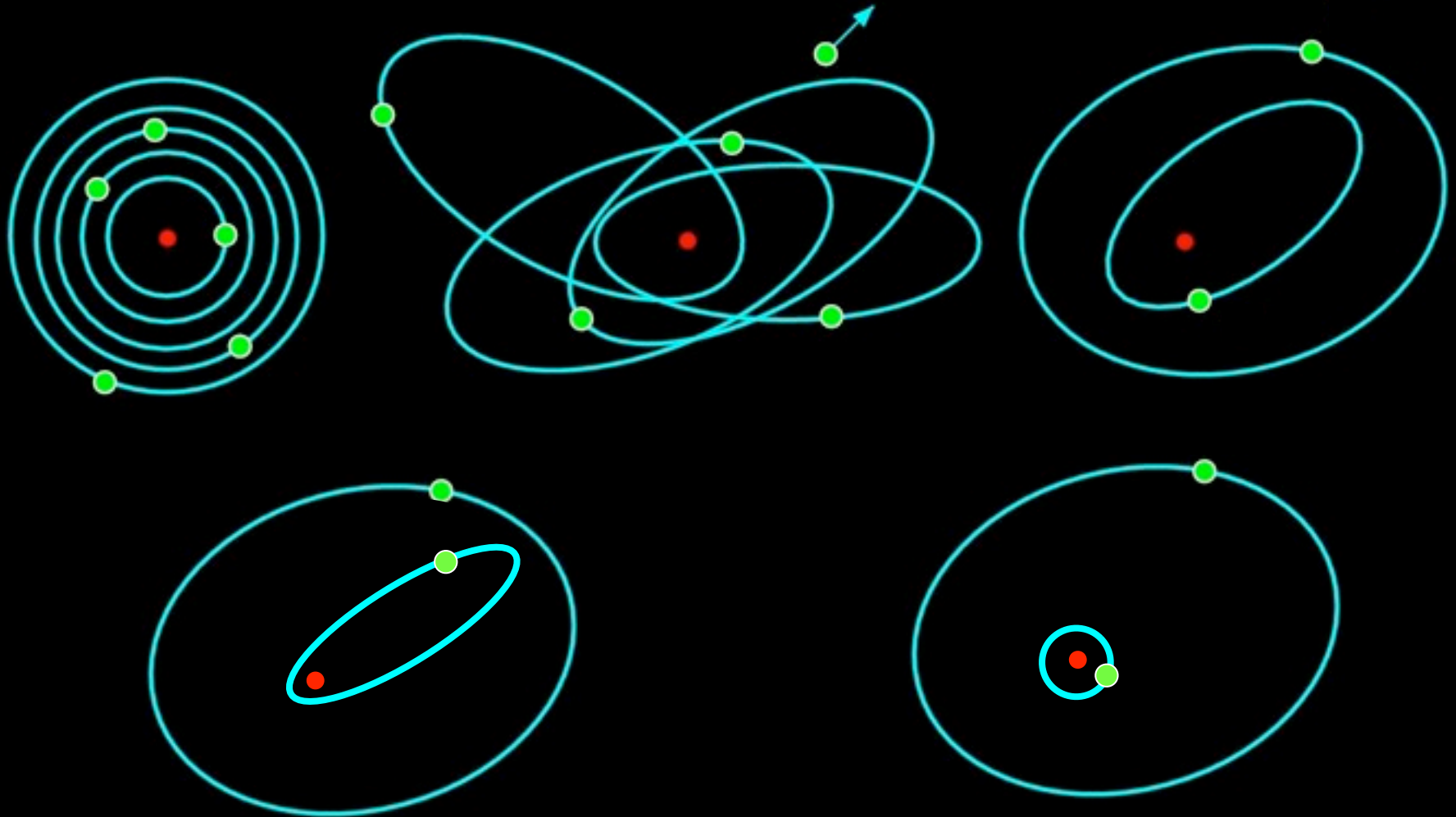
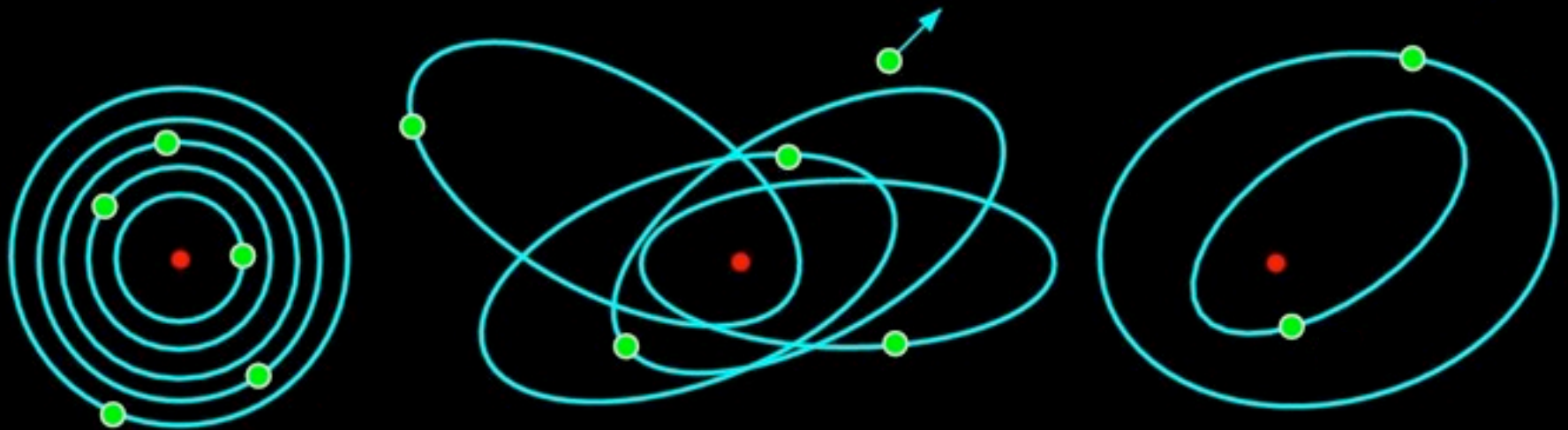


Illustration Adapted from E. Chiang

Hot Jupiters via Planet Scattering + Tidal Circularization

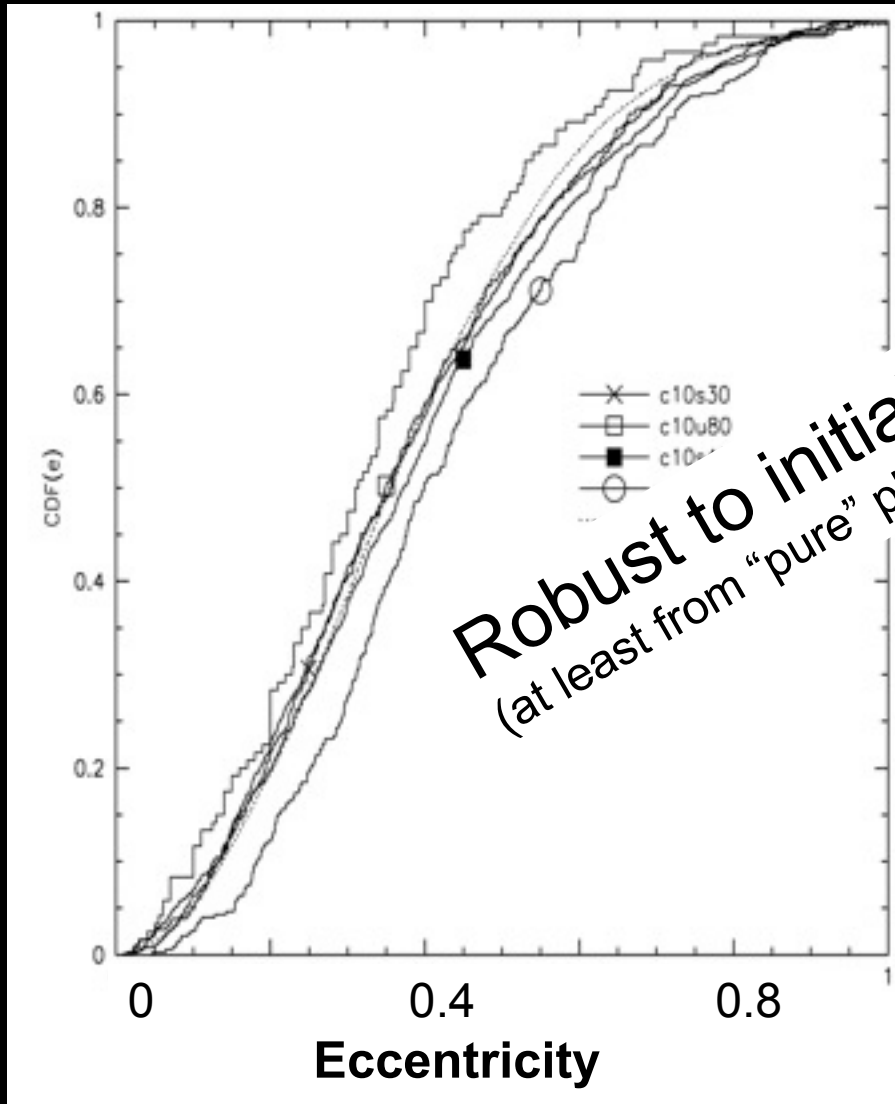


Eccentric Giant Planets via Planet Scattering



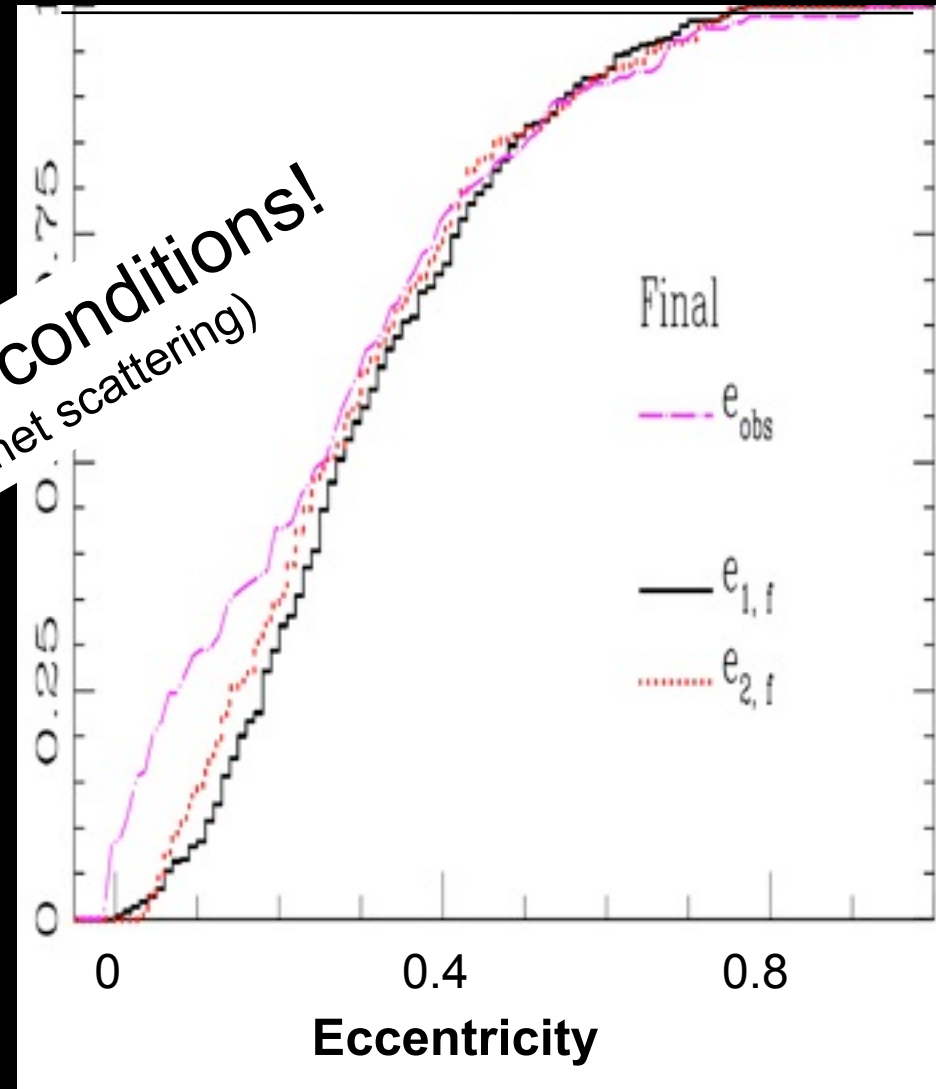
Eccentricity Distribution Predicted by Planet Scattering

Many Planets



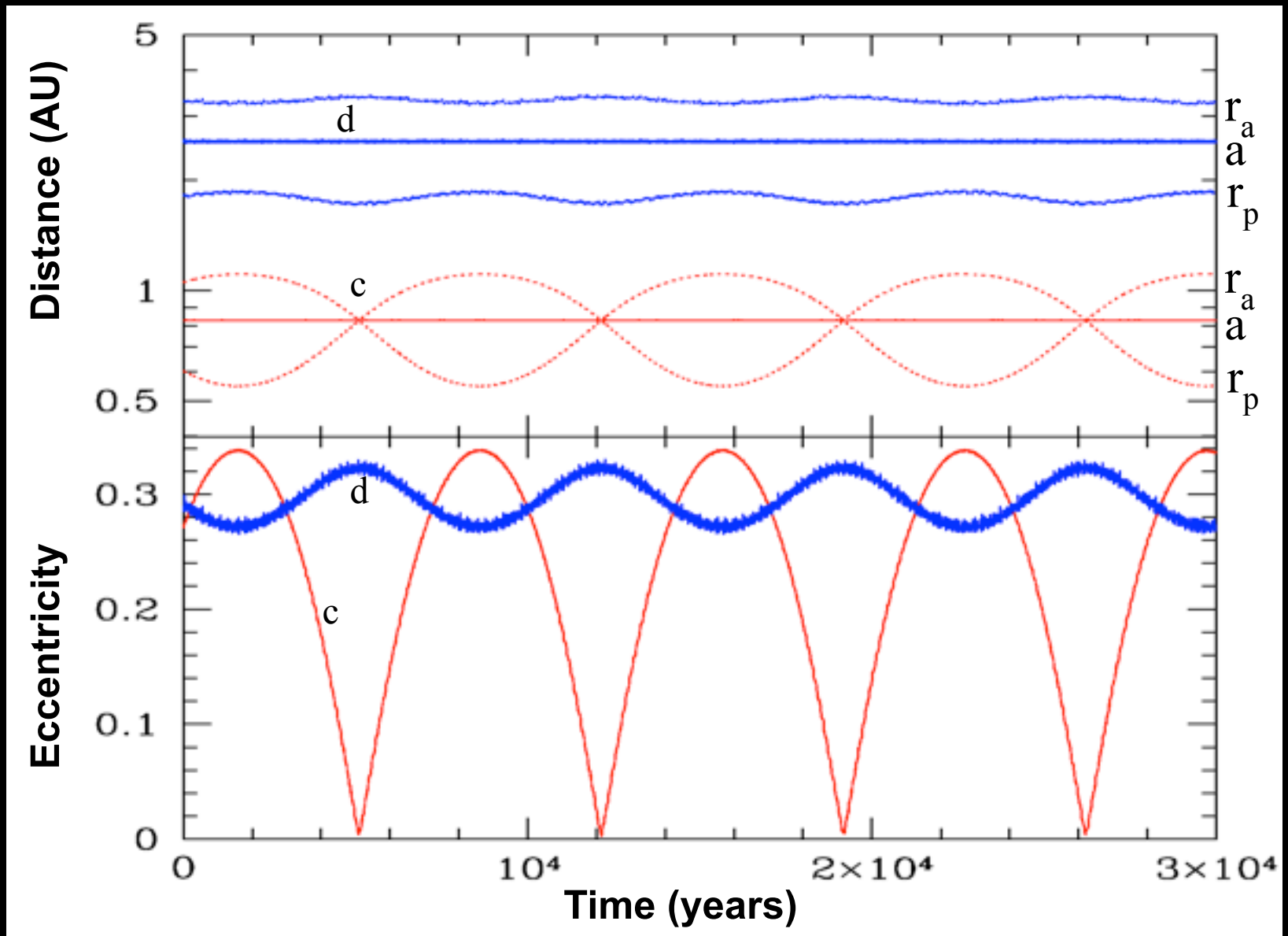
Juric & Tremaine 2007

Three Planets



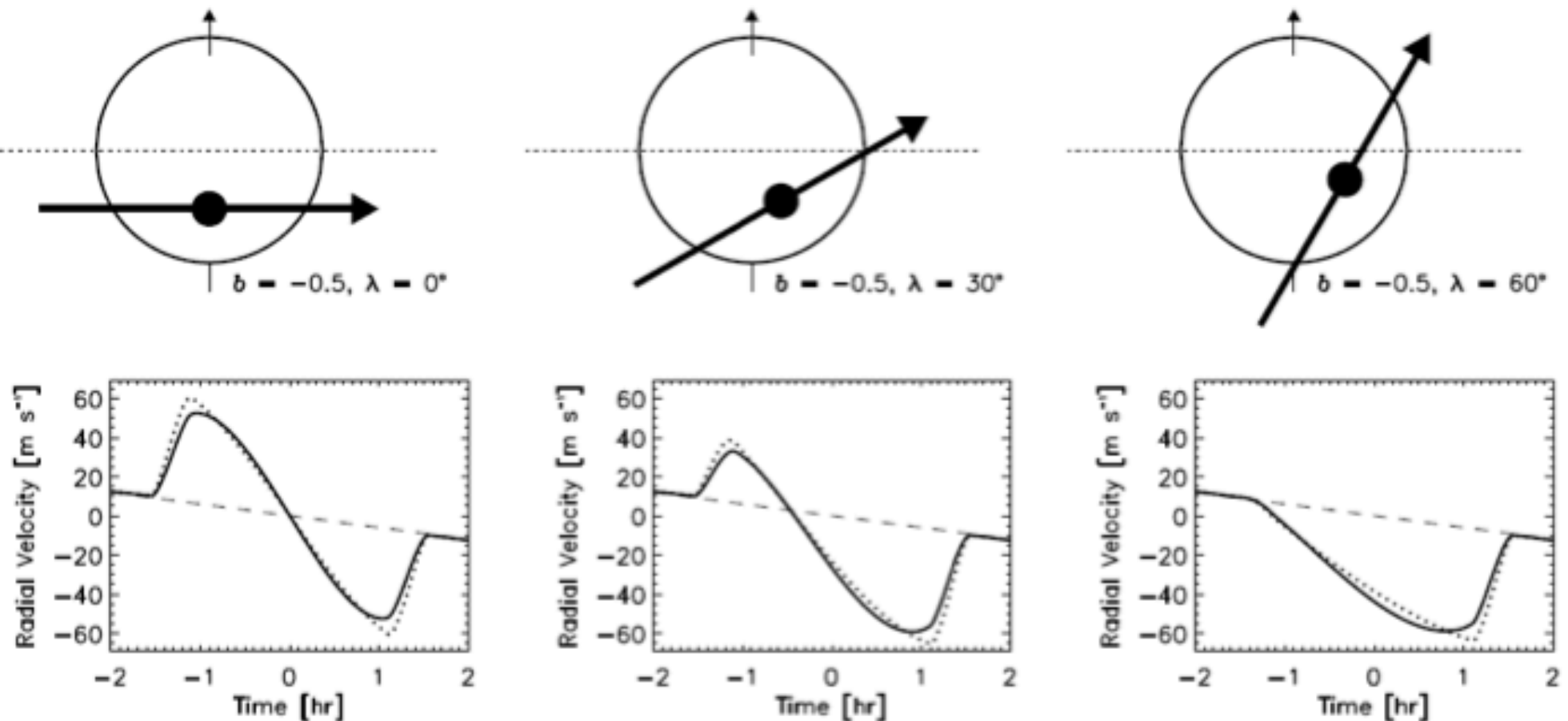
Chatterjee et al. 2007

Secular Evolution of Ups And

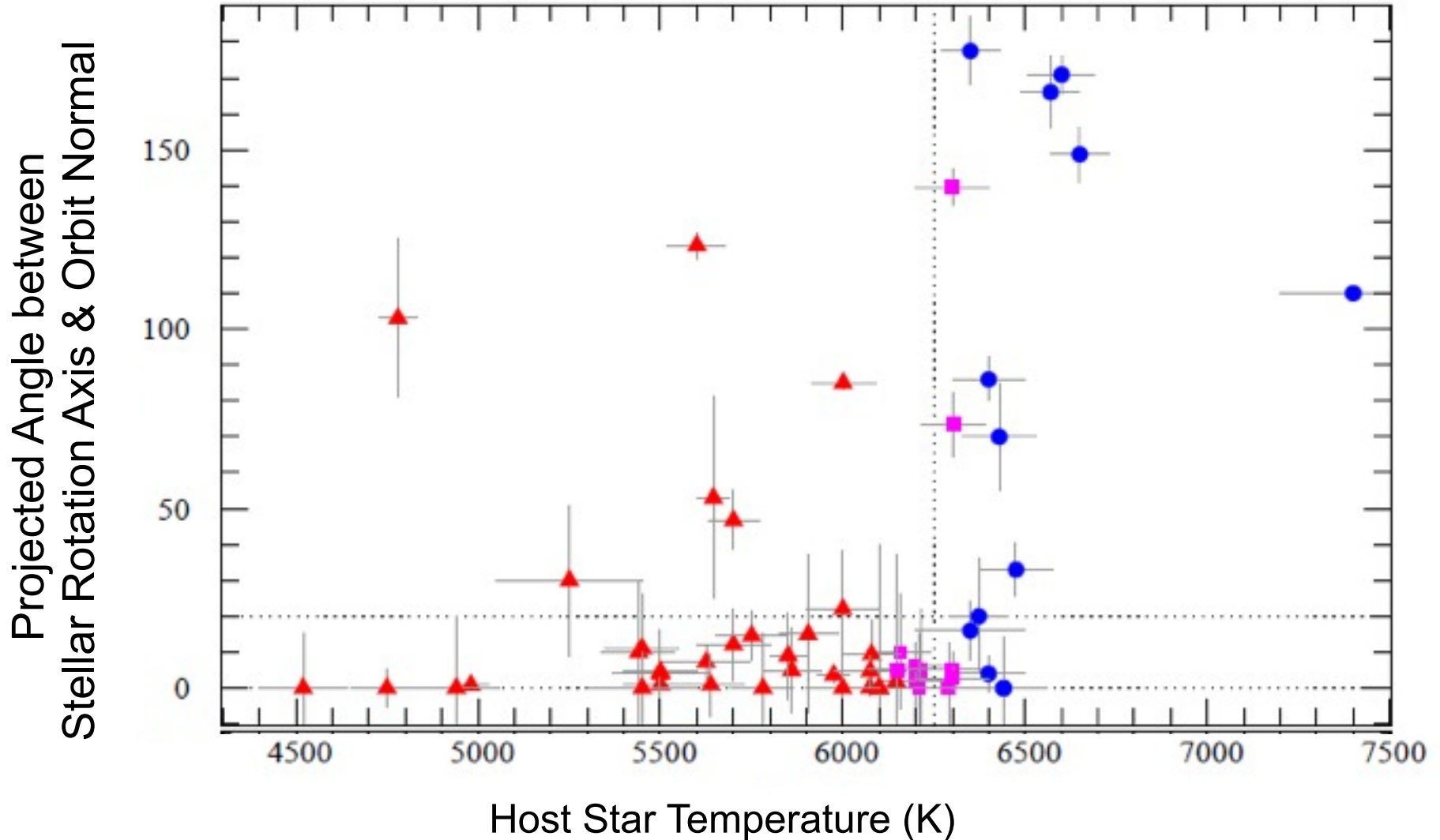


Measuring Exoplanet Inclinations

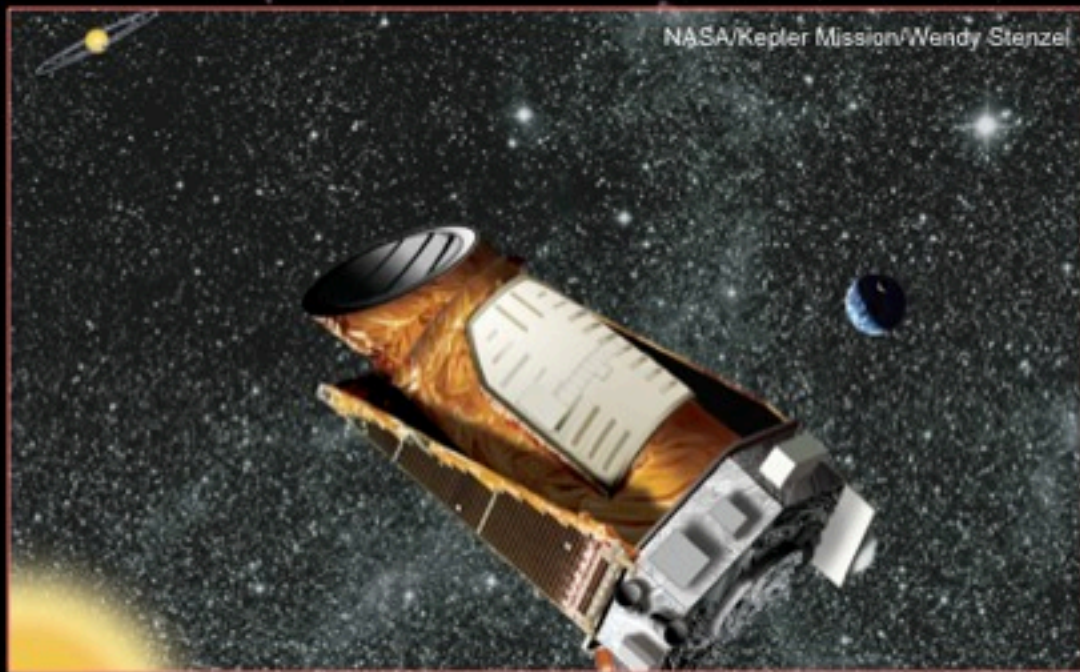
- Tidal dissipation in the planet rapidly damps eccentricity
- Search for planets with inclination excited by strong scattering (Chatterjee et al. 2008; Fabrycky & Tremaine 2007; Nagasawa et al. 2007)



Stars & Hot-Jupiter's can be Misaligned



Launch of *Kepler* Mission



Frequency of Earth-like planets

Kepler
Solar Orbit

1543

1609

1781

1846

1993

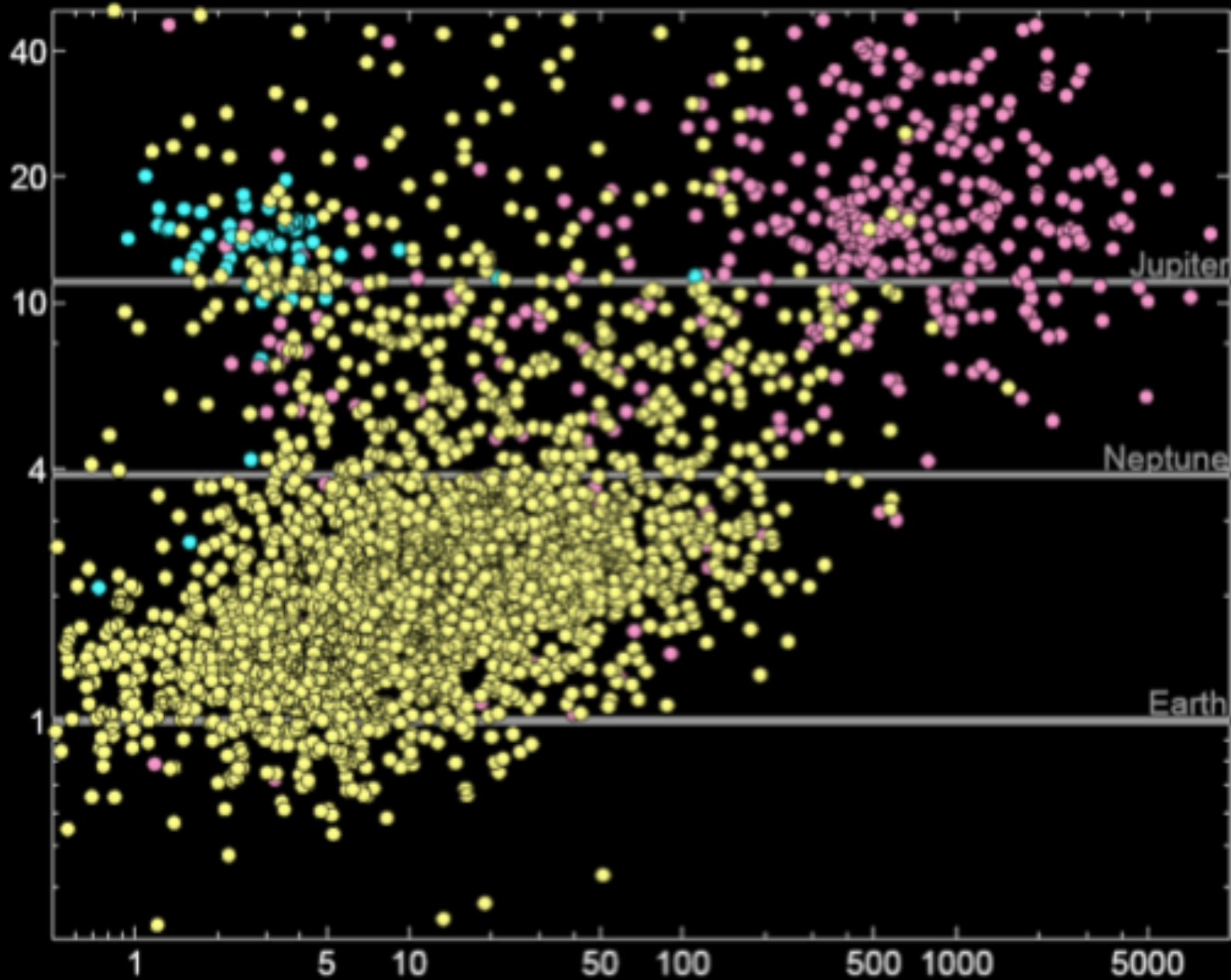
1995

1999

2001

2009

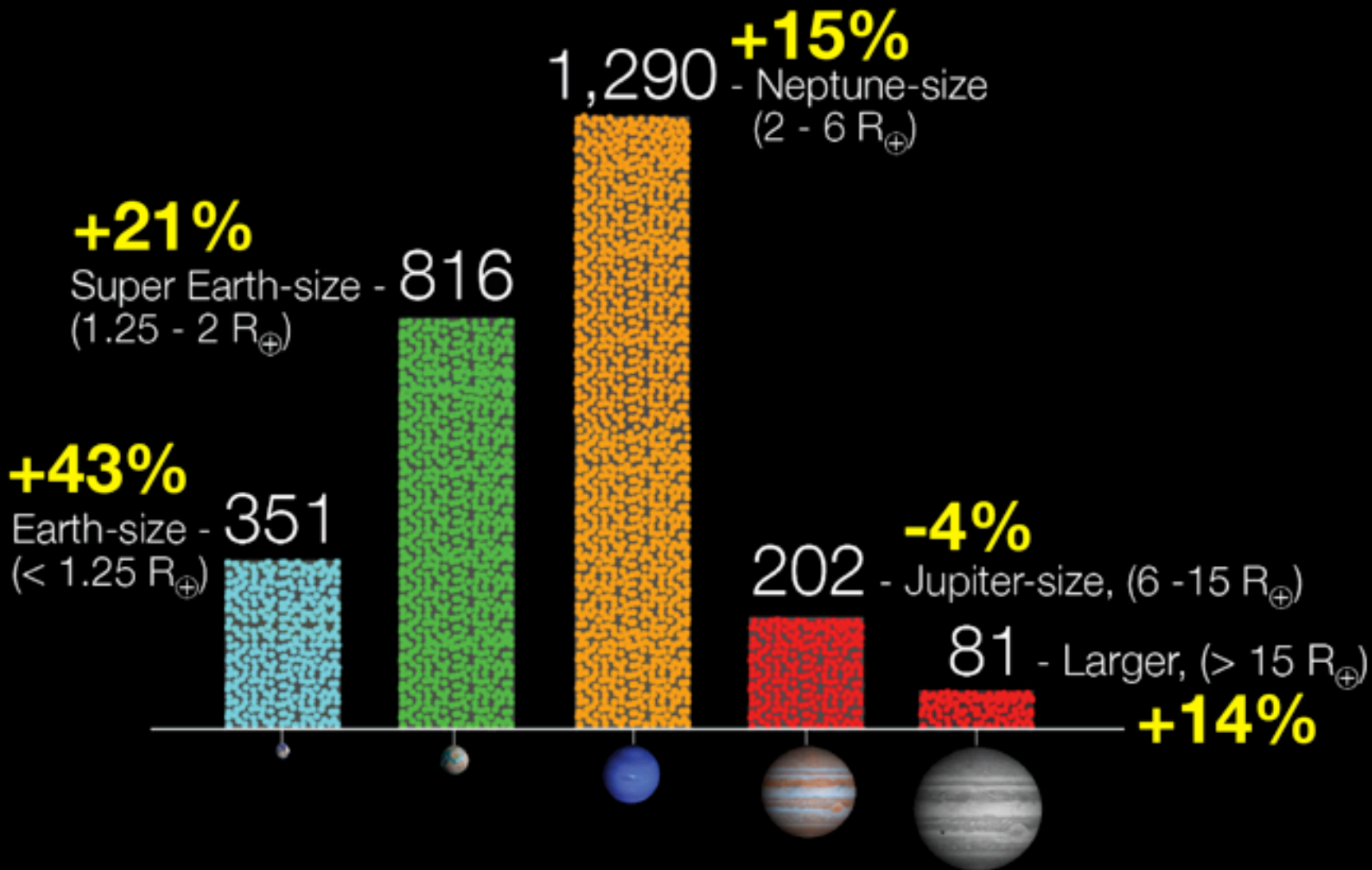
Planet Size (Earth Radii)



Orbital Period (days)

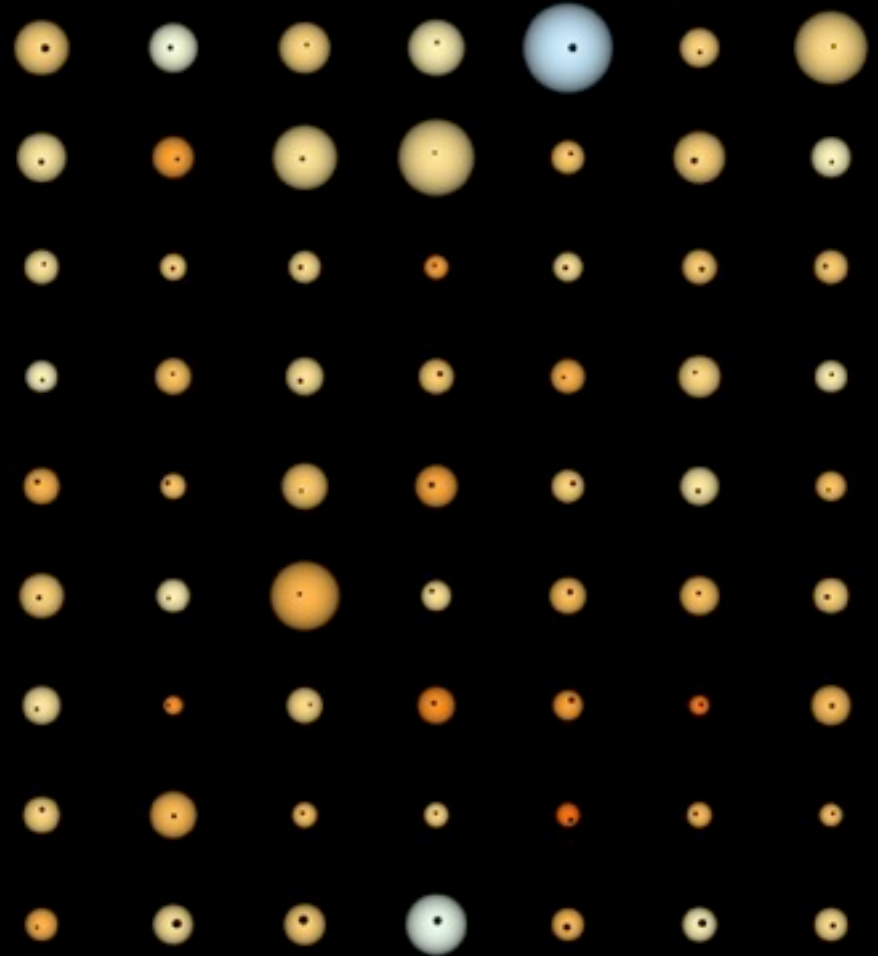
Sizes of Planet Candidates

As of January 7, 2013



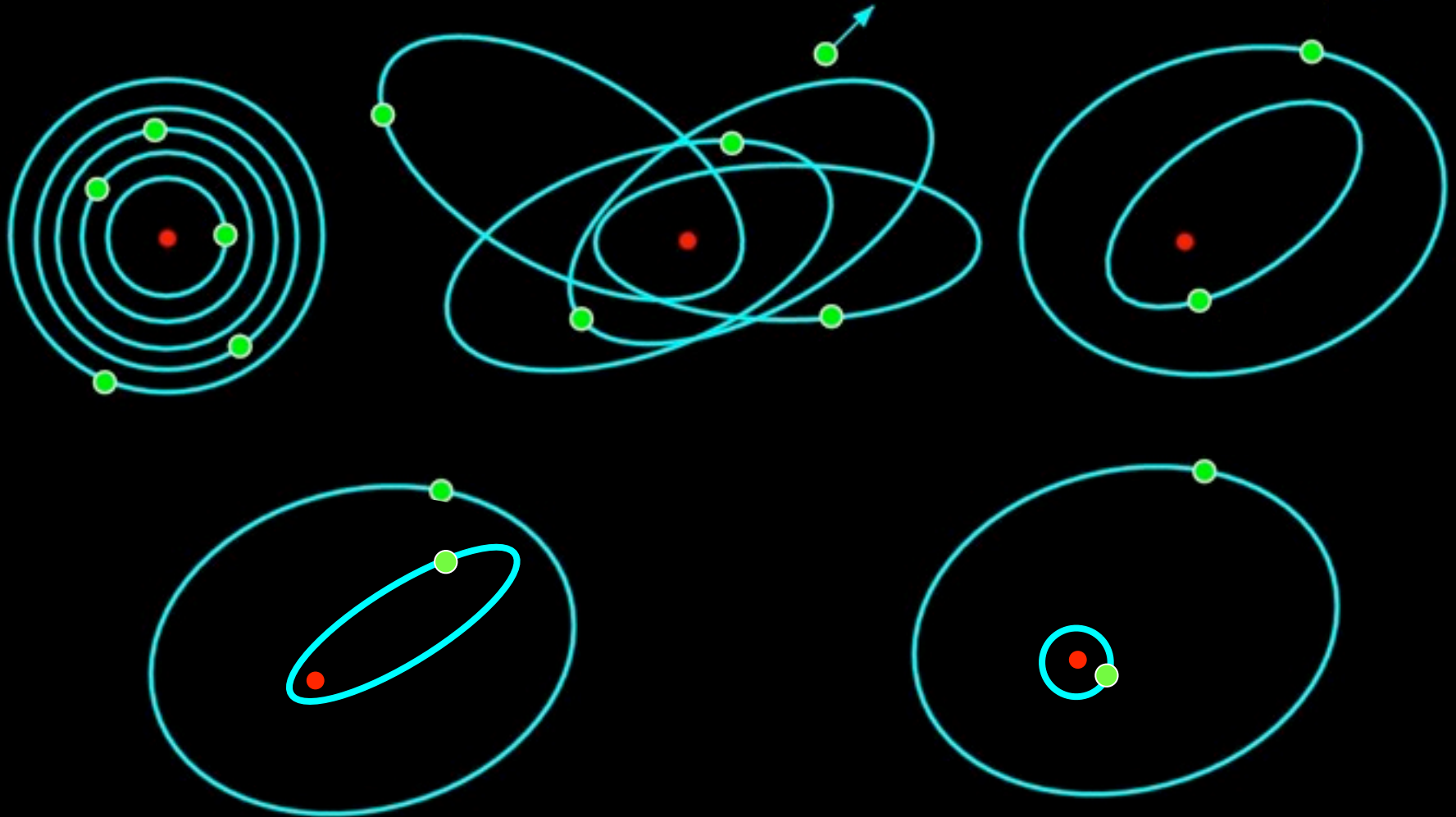
Hot Jupiters are Lonely

- 63 Hot Jupiters
- No other transiting planets
- No TTV signals
- Consistent with eccentricity excitation followed by tidal circularization

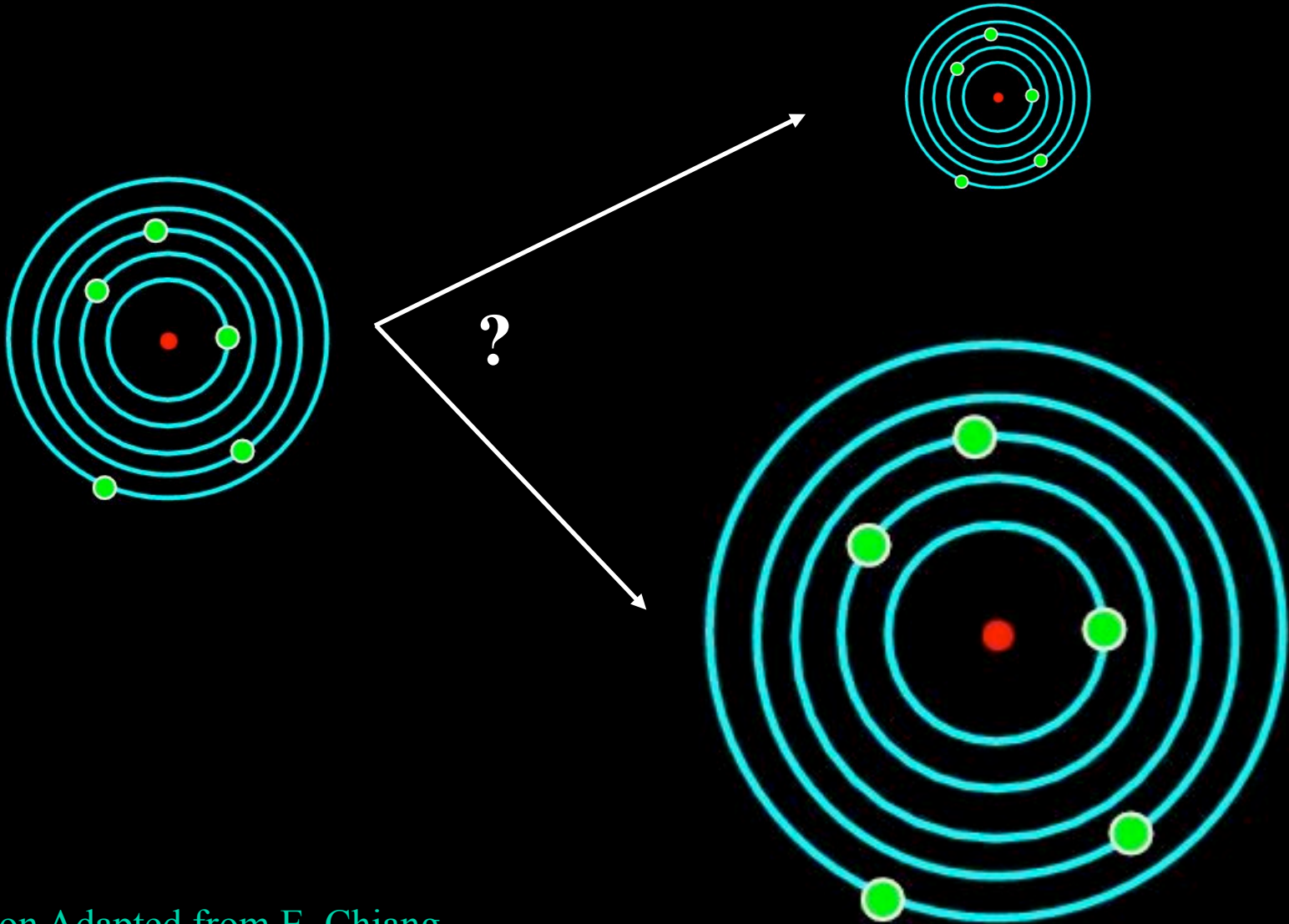


(Steffen et al. 2012 PNAS;
see Szabo et al. arxiv)

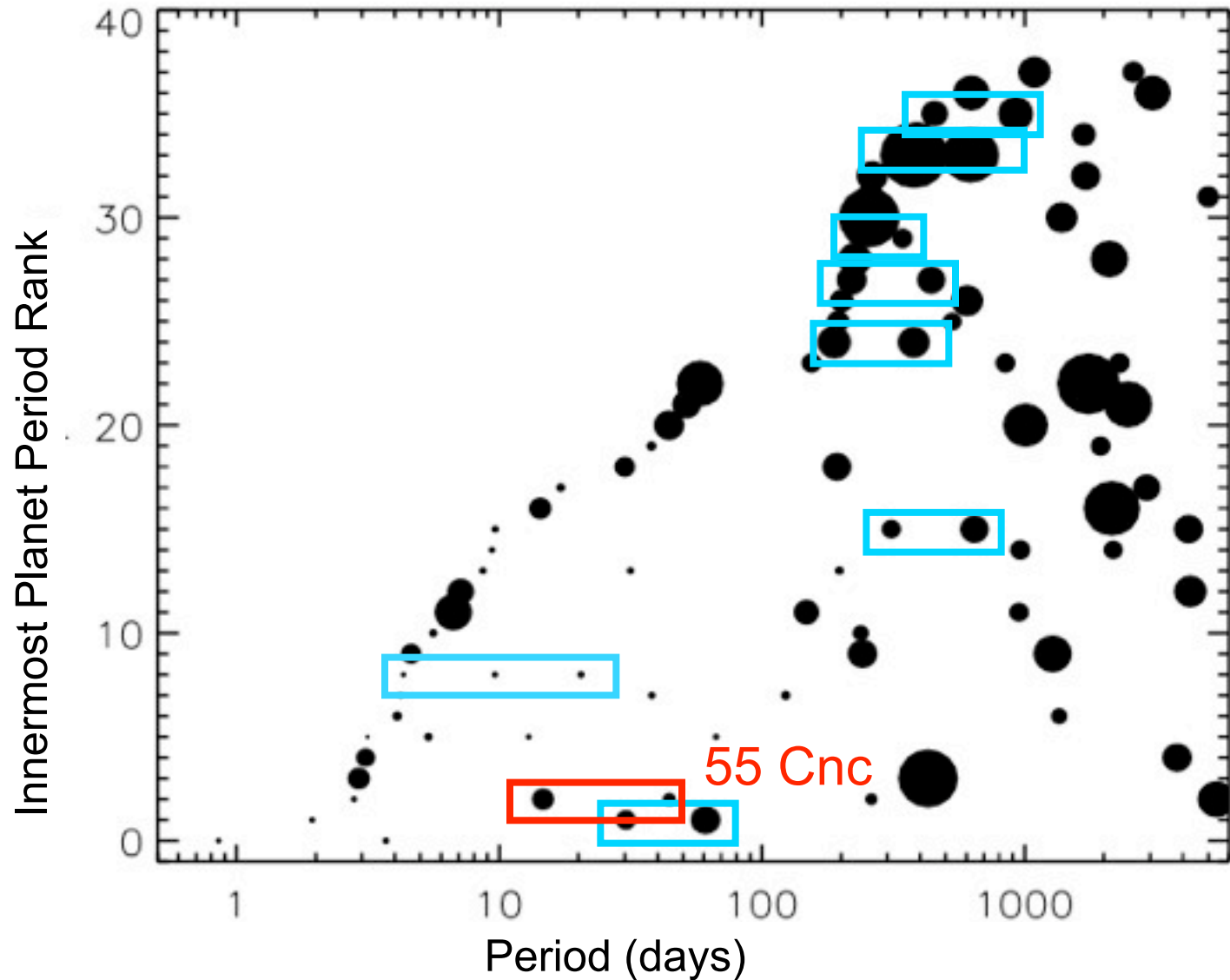
Hot Jupiters via Planet Scattering + Tidal Circularization



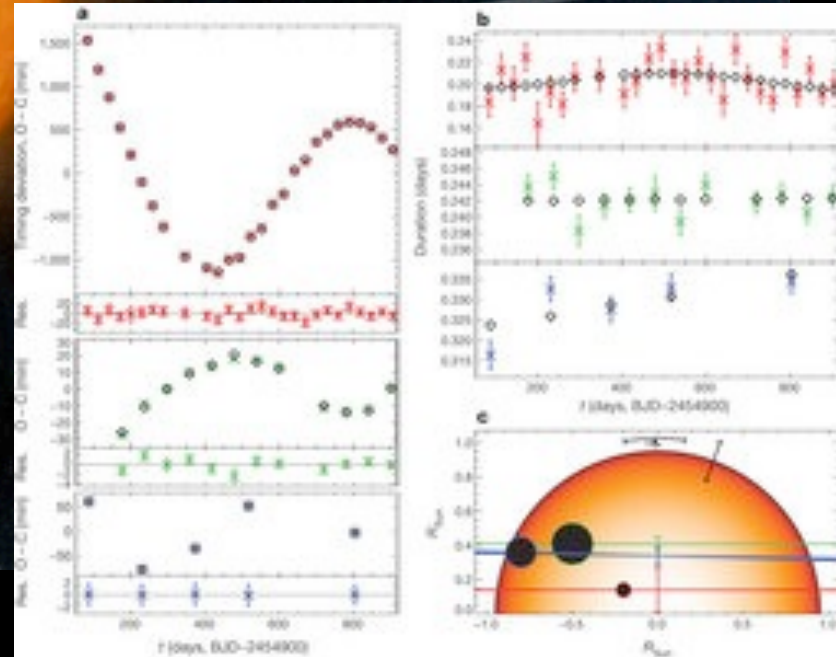
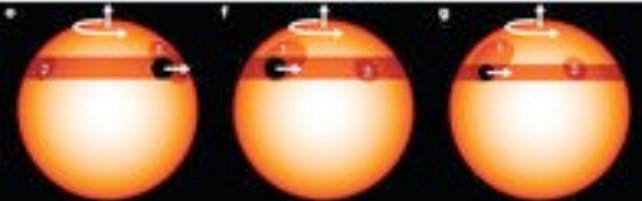
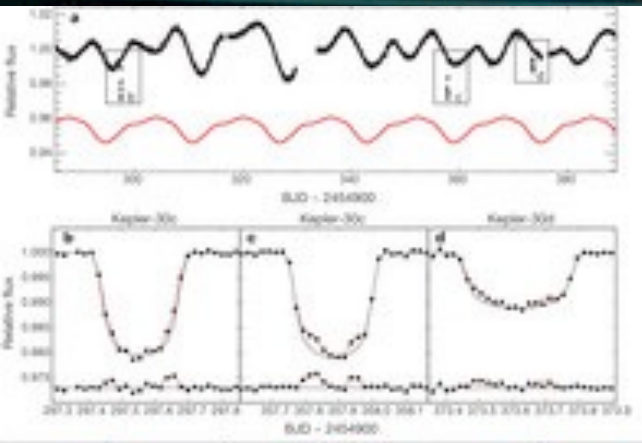
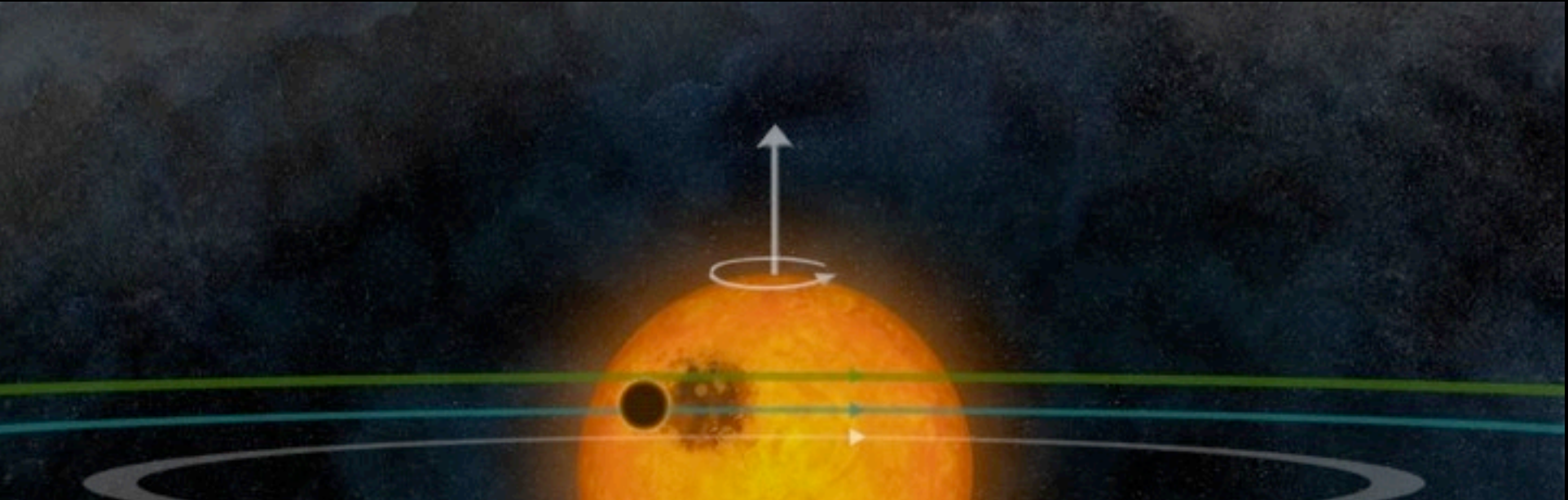
Hot Jupiters via Disk Migration?



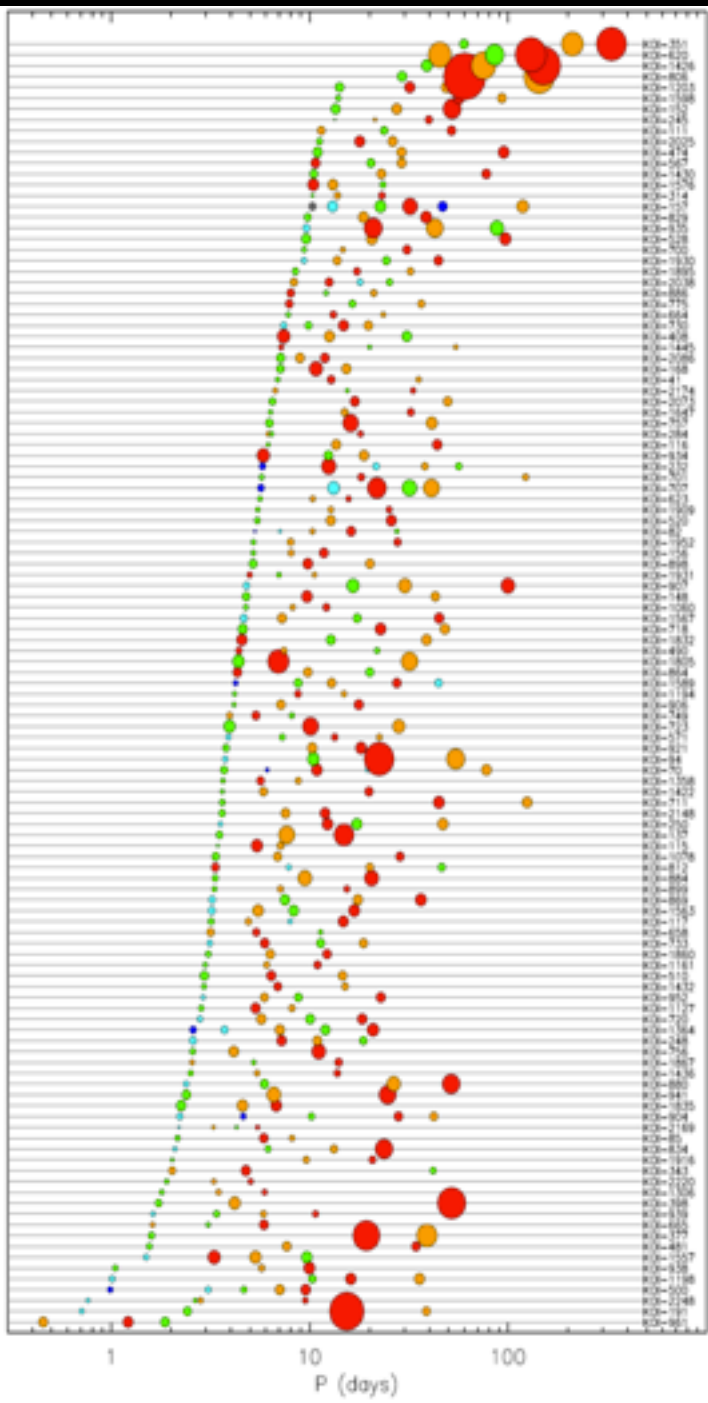
Orbital Resonances Among Multi-Planet Systems Discovered via RVs



Kepler-30: Coplanarity via Spot Crossings



Extremely Compact Multi-transiting Planetary Systems

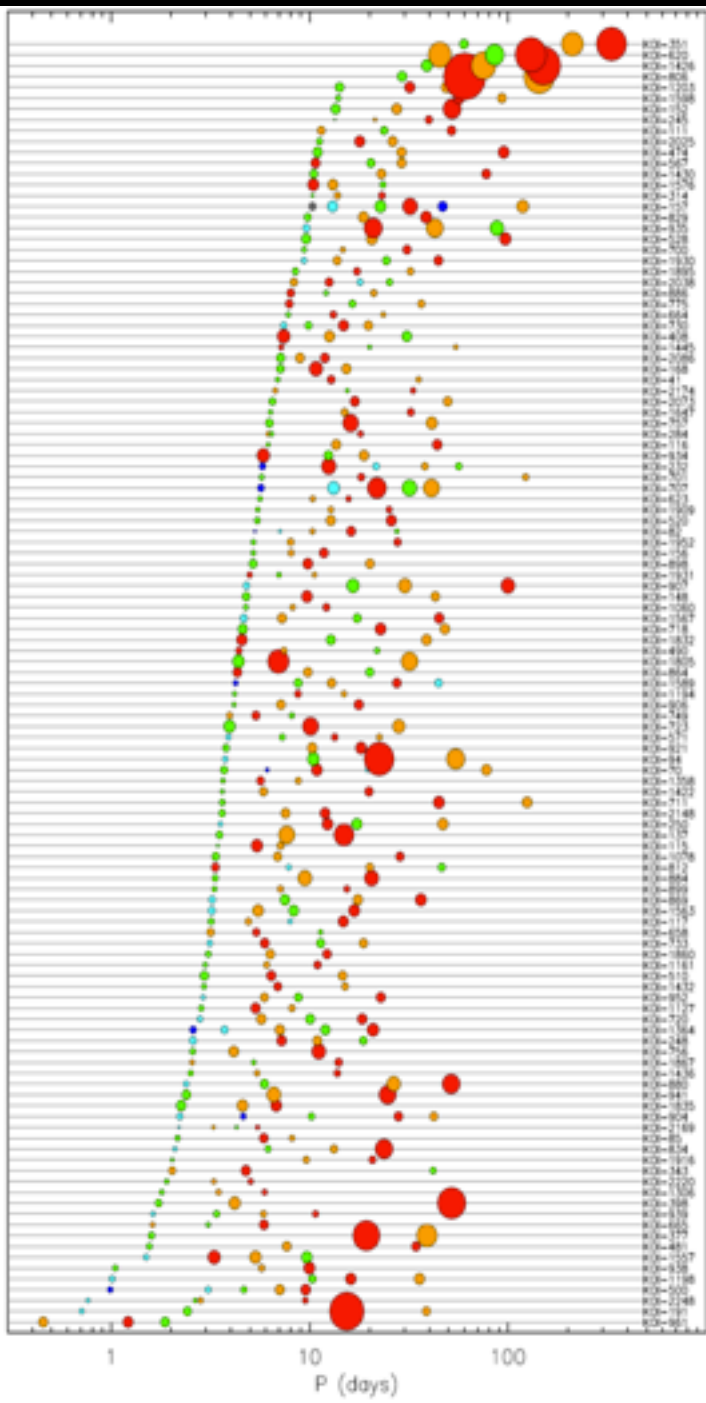


Fabrycky et al. 2012

Extremely Compact Multi-transiting Planetary Systems

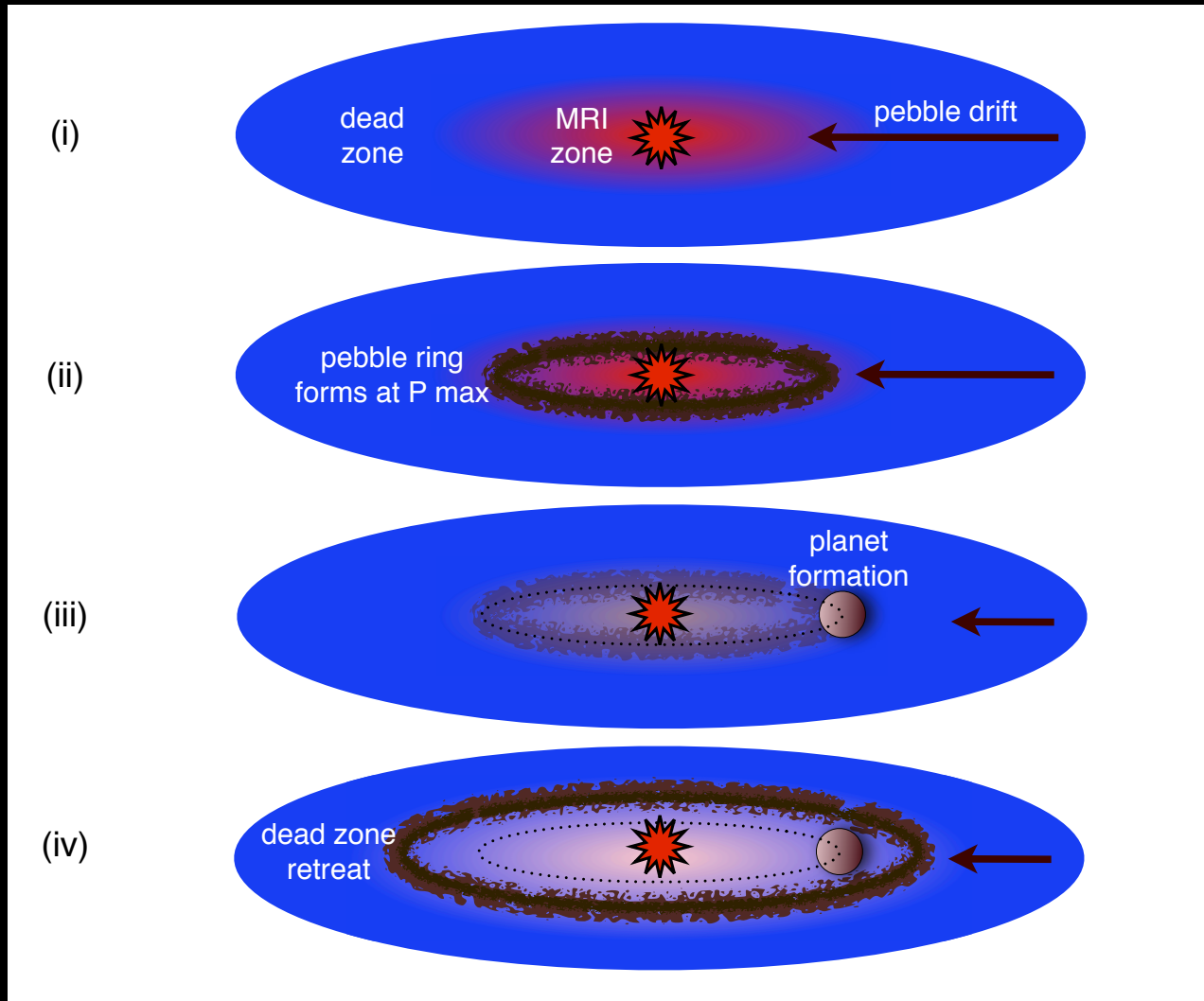
Higher solid density close to star-
idea of minimum mass extrasolar nebula
(Laughlin et al. 2012, also see Hansen &
Murray 2012)

Inside-out planet formation
(Chatterjee & Tan)



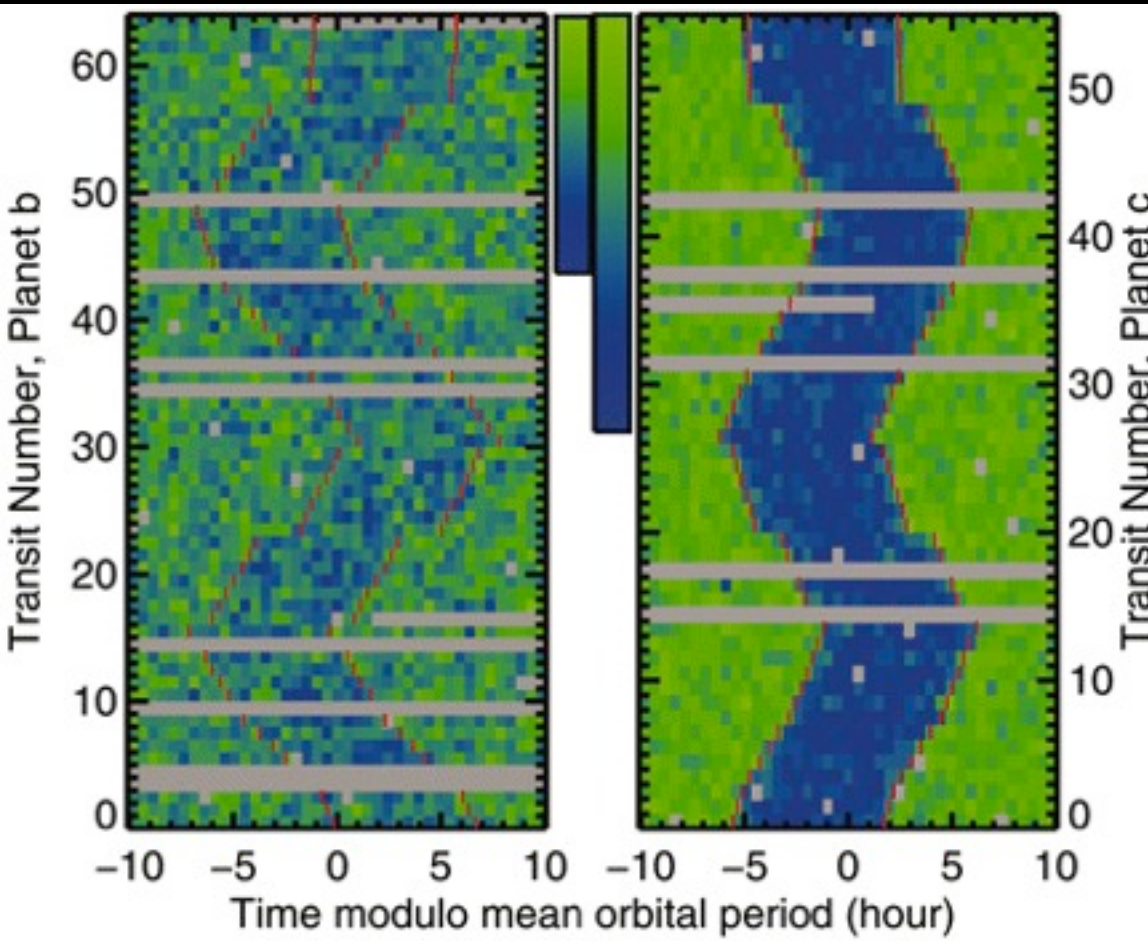
Fabrycky et al. 2012

Extremely Compact Multi-transiting Planetary Systems



Very Tightly Packed Planetary Systems

Kepler-36b&c: Chaotic due to 29:34 and 6:7 MMRs!

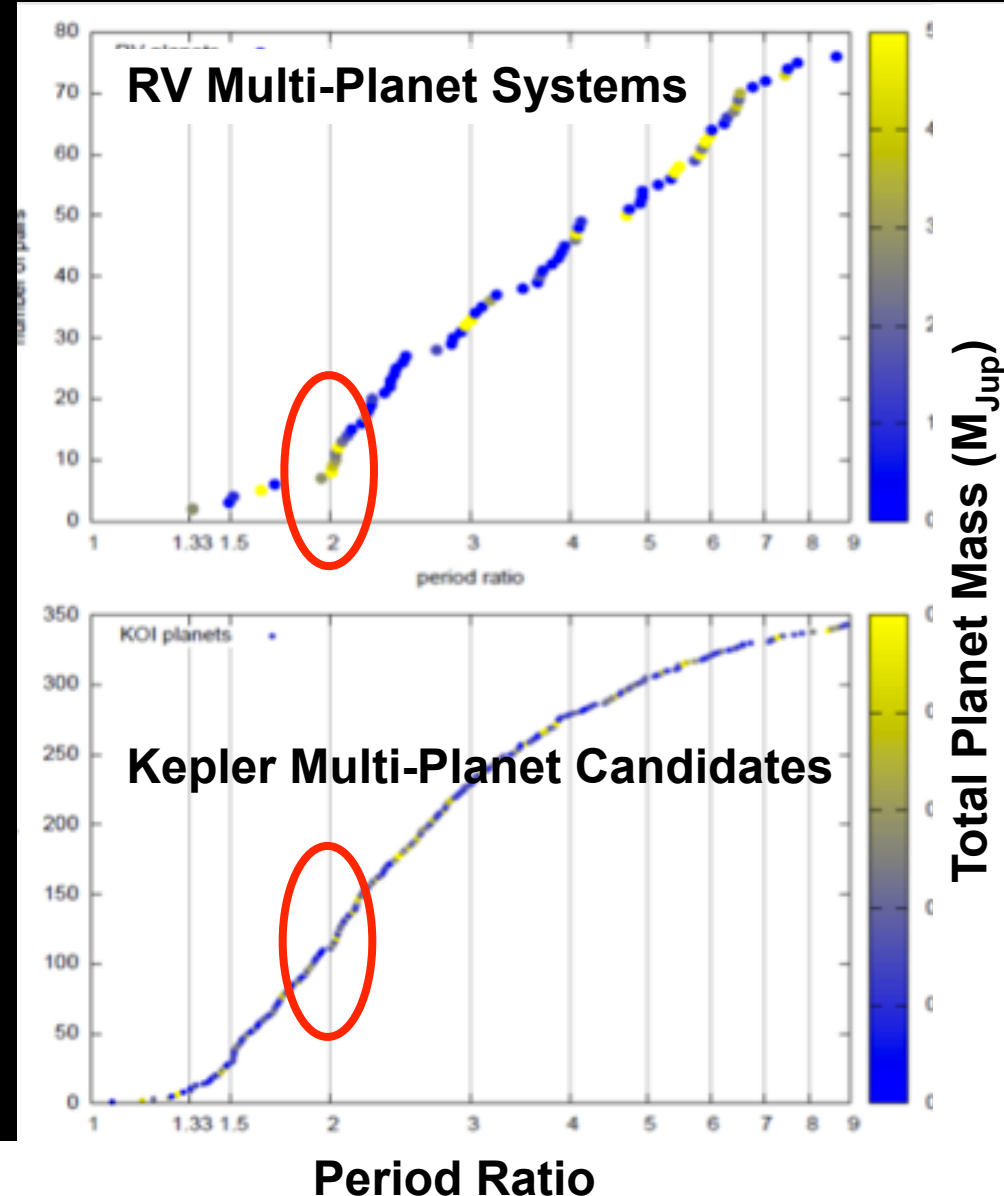


Carter et al. 2012; Deck et al. 2012

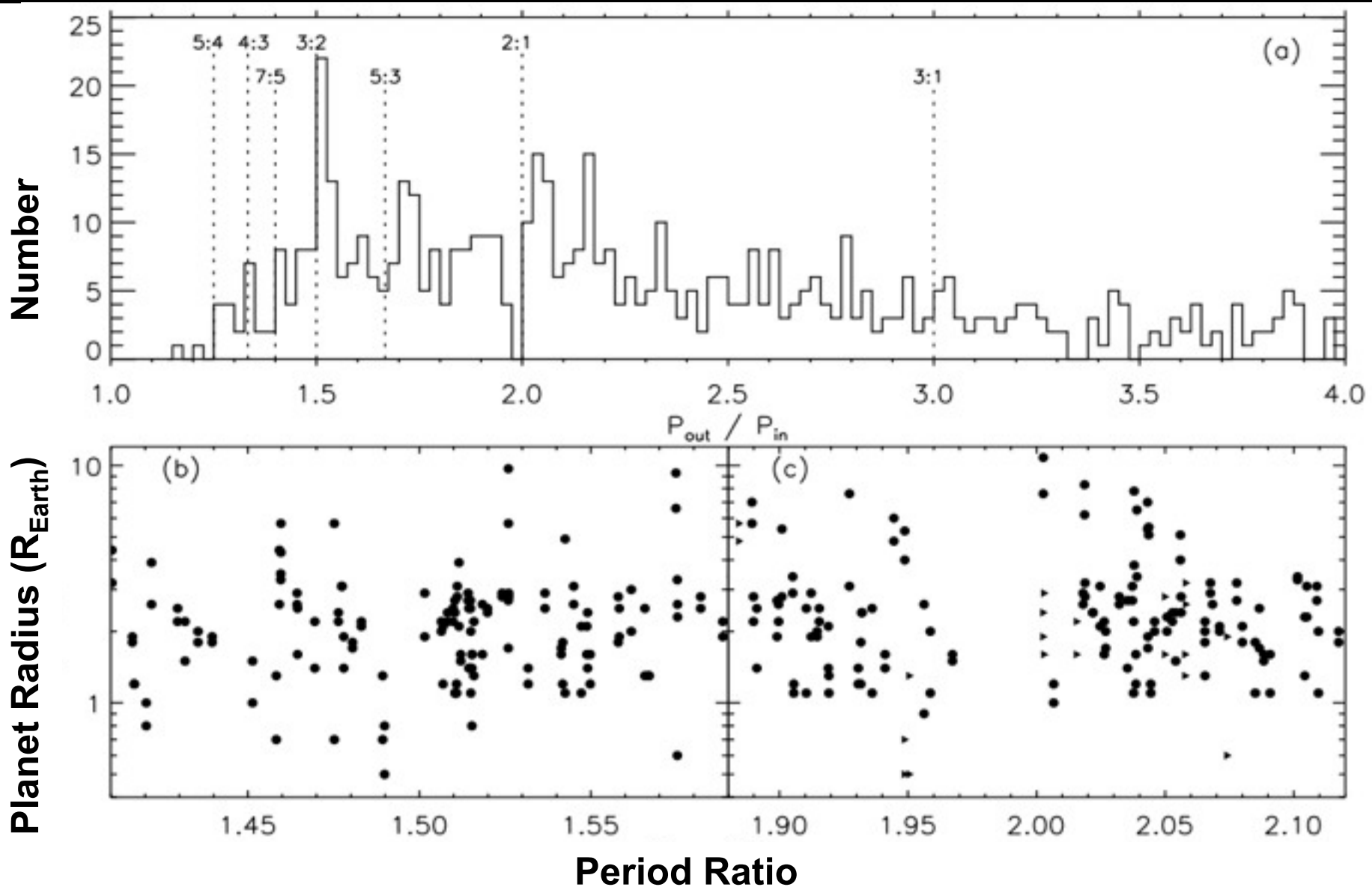
Resonances in Kepler Multi-Planet Systems

- Rarer than in RV systems
 - Predicted!
- Most near, but not in resonance
- Near resonant great for TTVs
 - esp. closely spaced pairs!

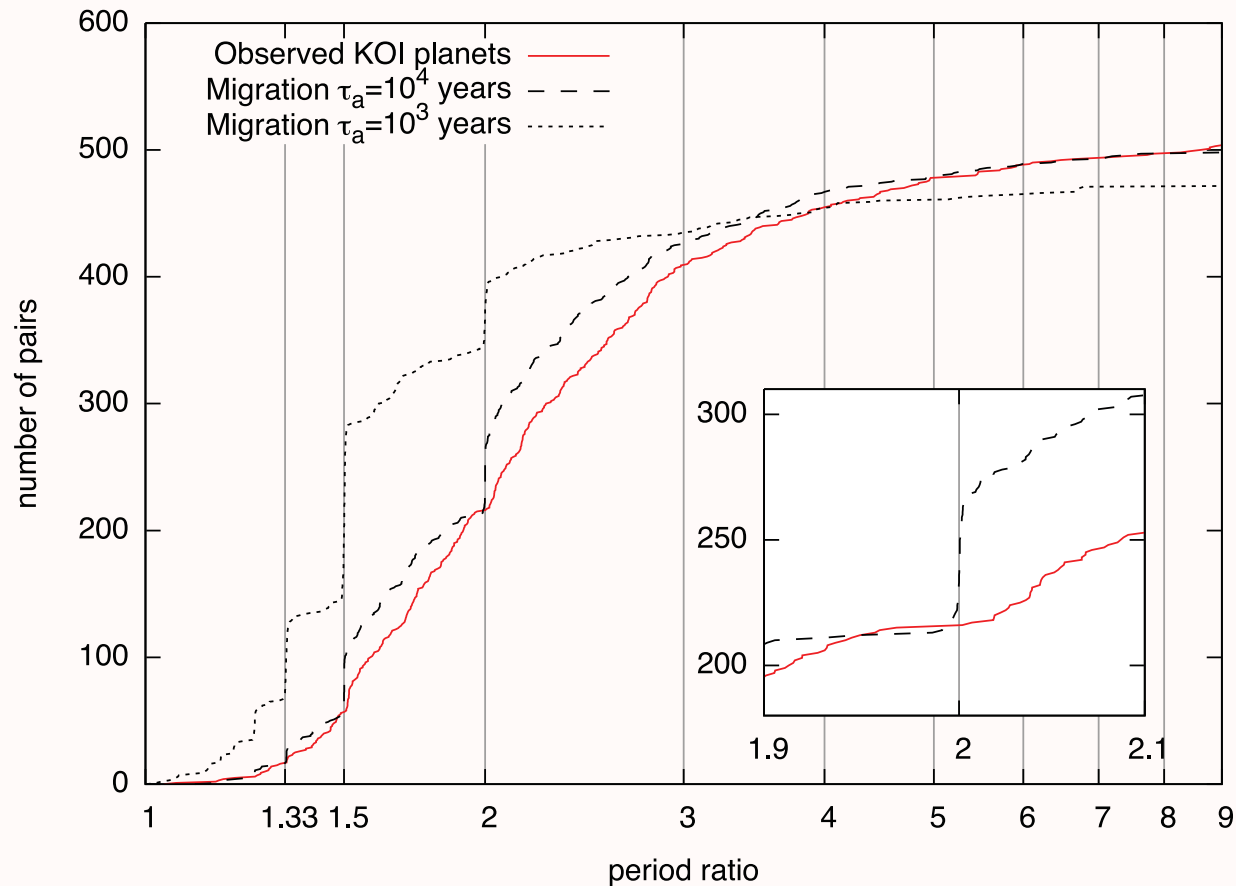
Cumulative Number of Planet Pairs



Kepler's Near Resonant Systems

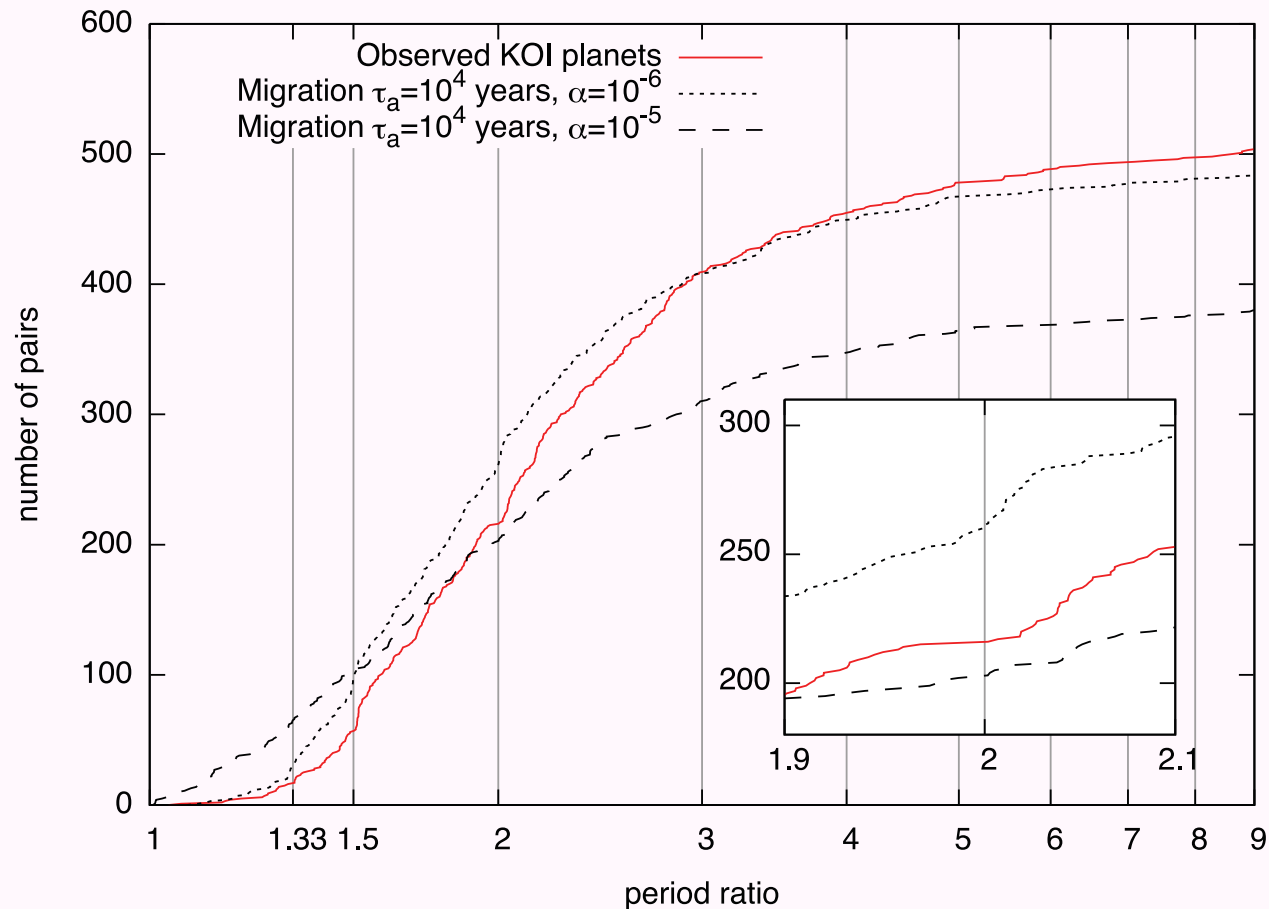


Kepler's Near Resonant Systems



Rein et al. 2012; see also Ford & Rasio 2008; Veras et al. 2012

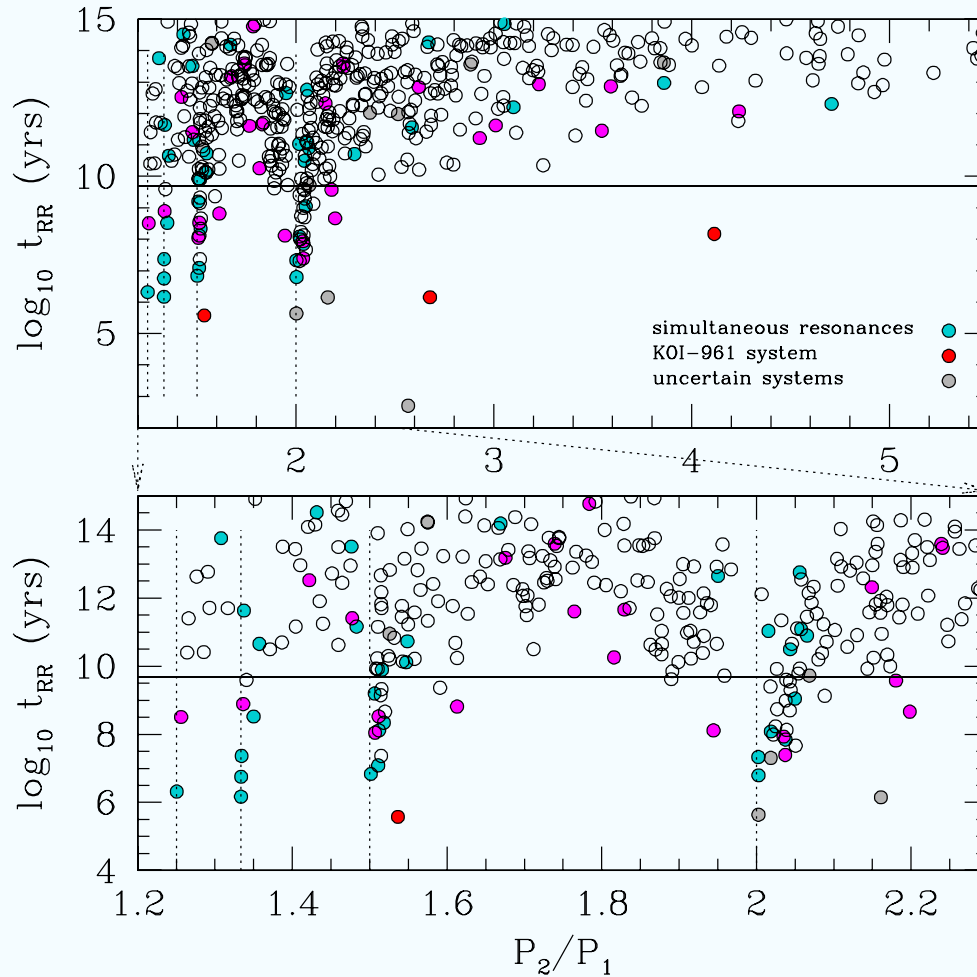
Kepler's Near Resonant Systems



Rein et al. 2012; see also Ford & Rasio 2008; Veras et al. 2012

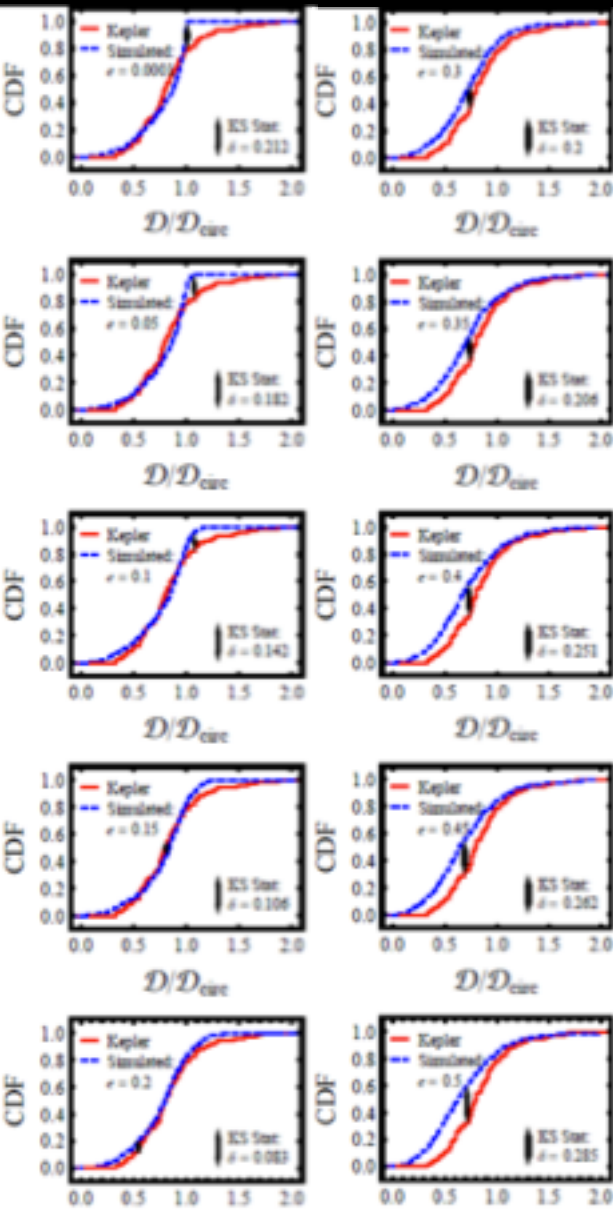
Kepler's Near Resonant Systems

resonant repulsion

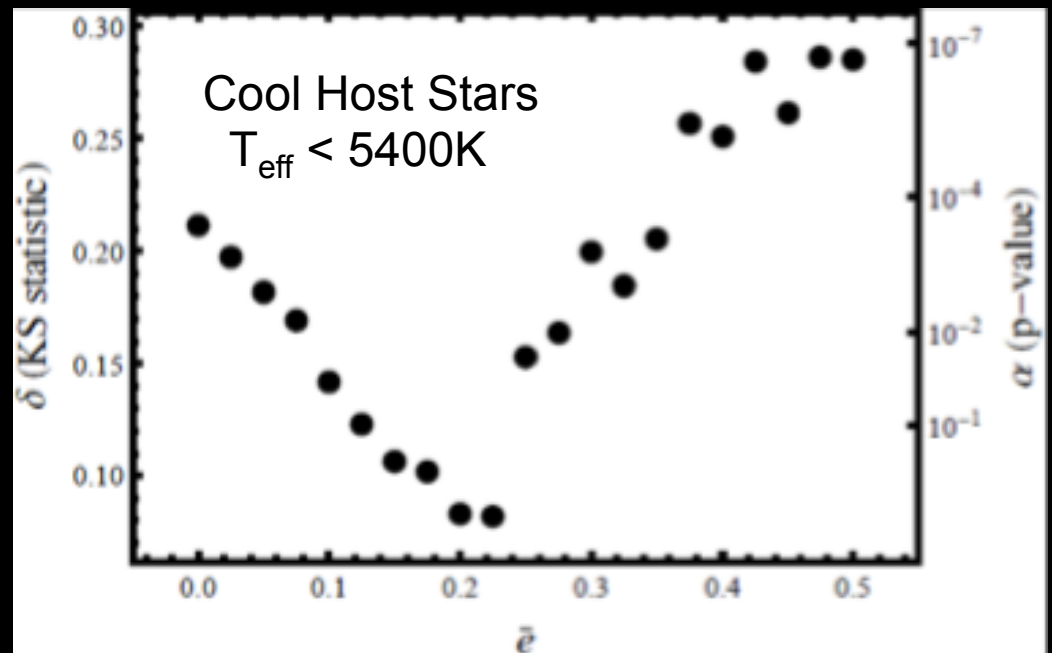


Lithwick et al. 2012; see also Ford & Rasio 2008; Veras et al. 2012

Eccentricities of Transiting Planets via Transit Duration Distribution



- Consistent w/ RV distribution
- Smaller planets have smaller eccentricities
- Subject to uncertainties in stellar properties (A. Moorhead+ 2011)



Testing Planet Formation Theory with Kepler

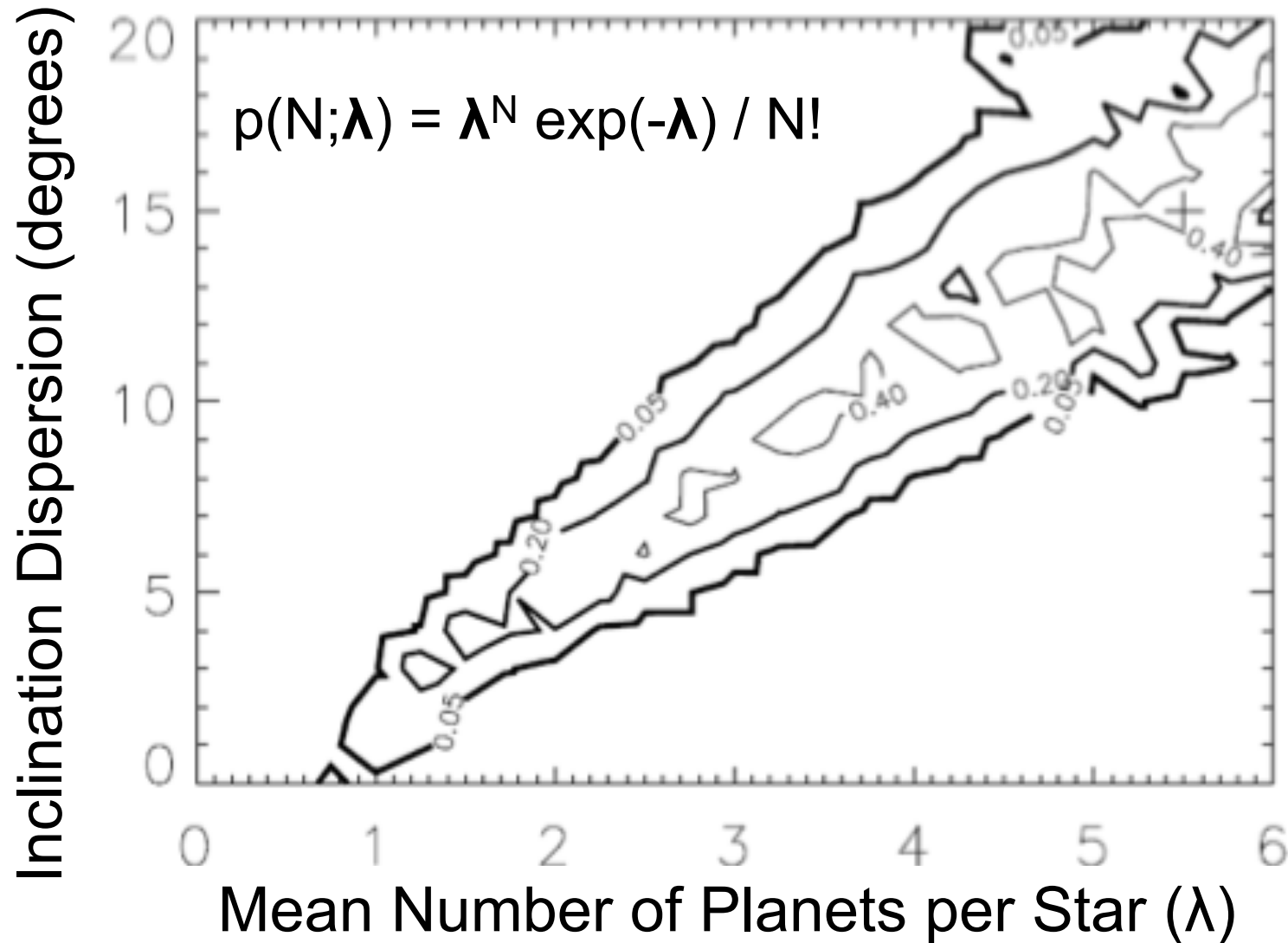
Orbital eccentricities, inclinations & multiplicity are key probes of planet formation:

- Eccentricity distribution (+ stellar densities) → Transit duration distribution
- **Inclination distribution + Frequency of multiple planet systems** (+ period distribution) → **Frequency of multiply transiting systems & transit duration variations**
- Frequency of multiple planet systems + Eccentricity distribution (+ period distribution) → Distribution of TTV signatures

One complex inverse problem!

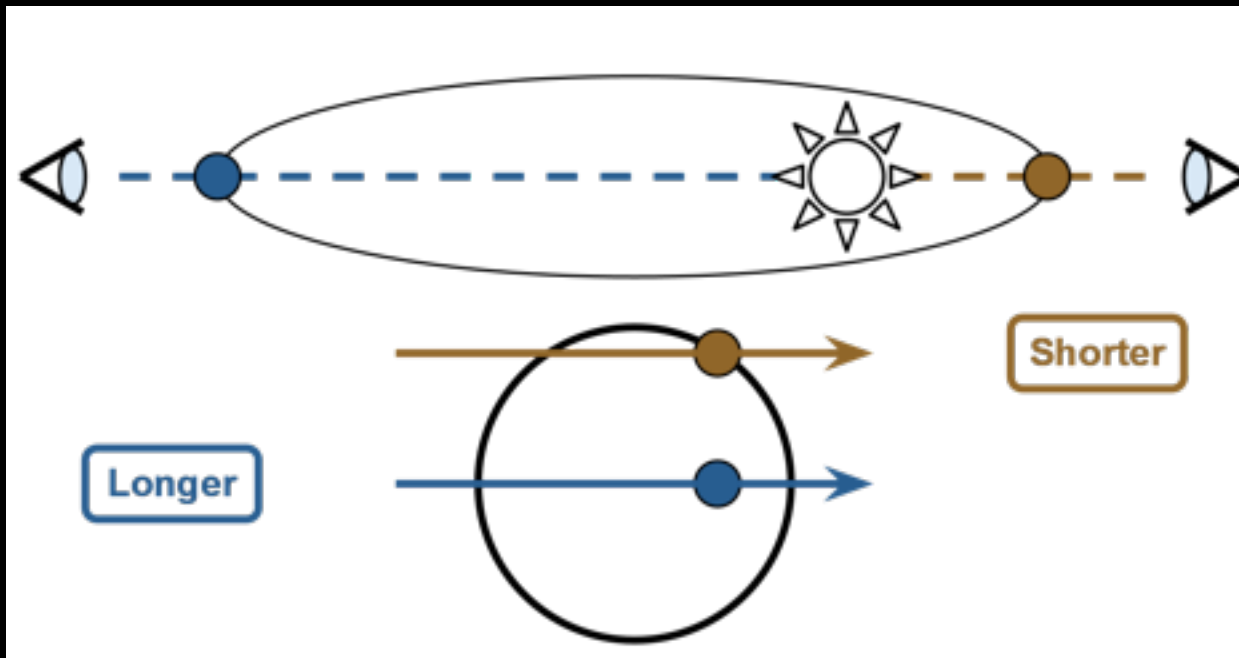
(Observables, Desired Distributions, Both)

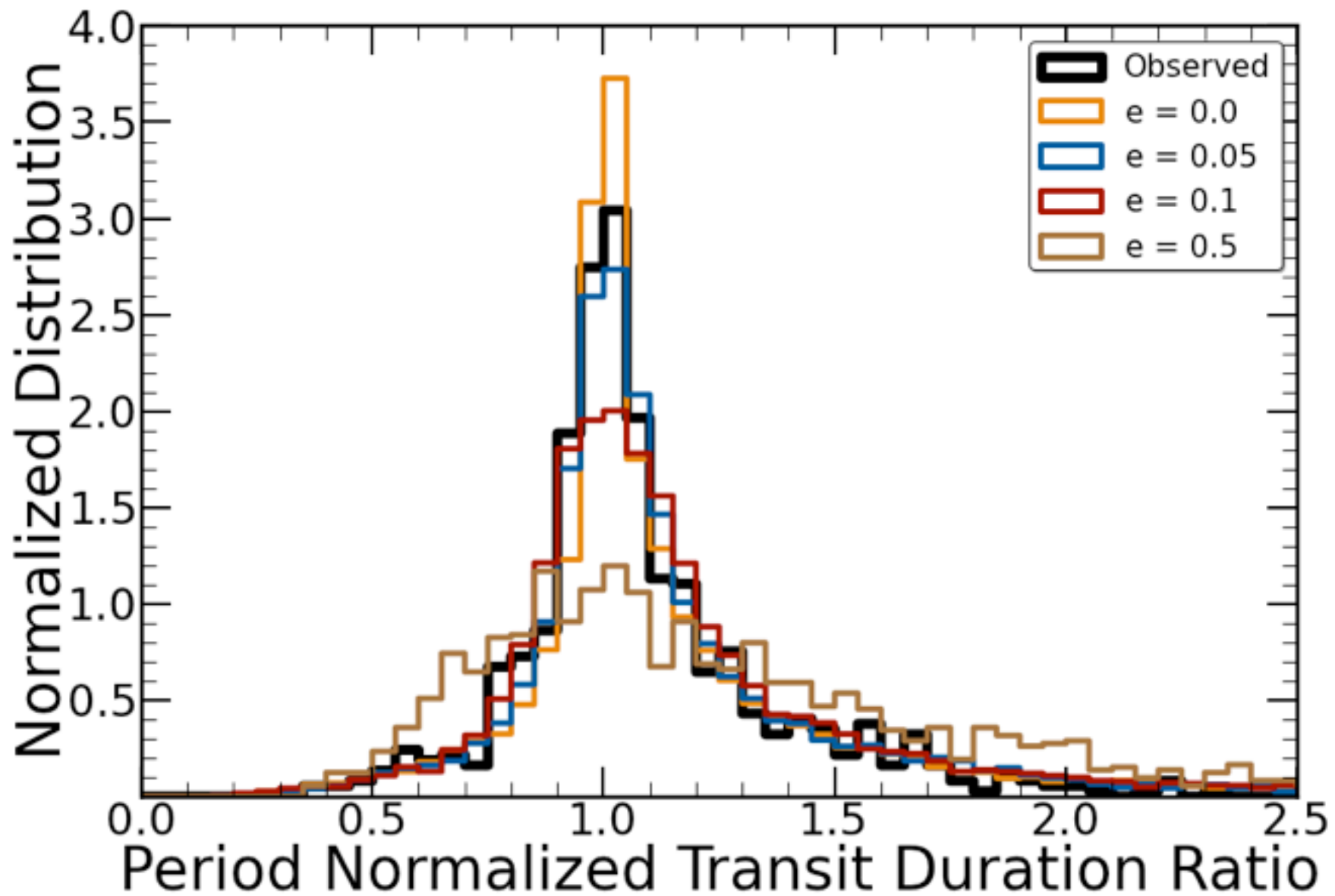
Planet Multiplicity & Mutual Inclinations



Period-Normalized Transit Duration Ratio

$$\xi \equiv \left(D_{in} \right)$$





Testing Planet Formation Theory with Kepler

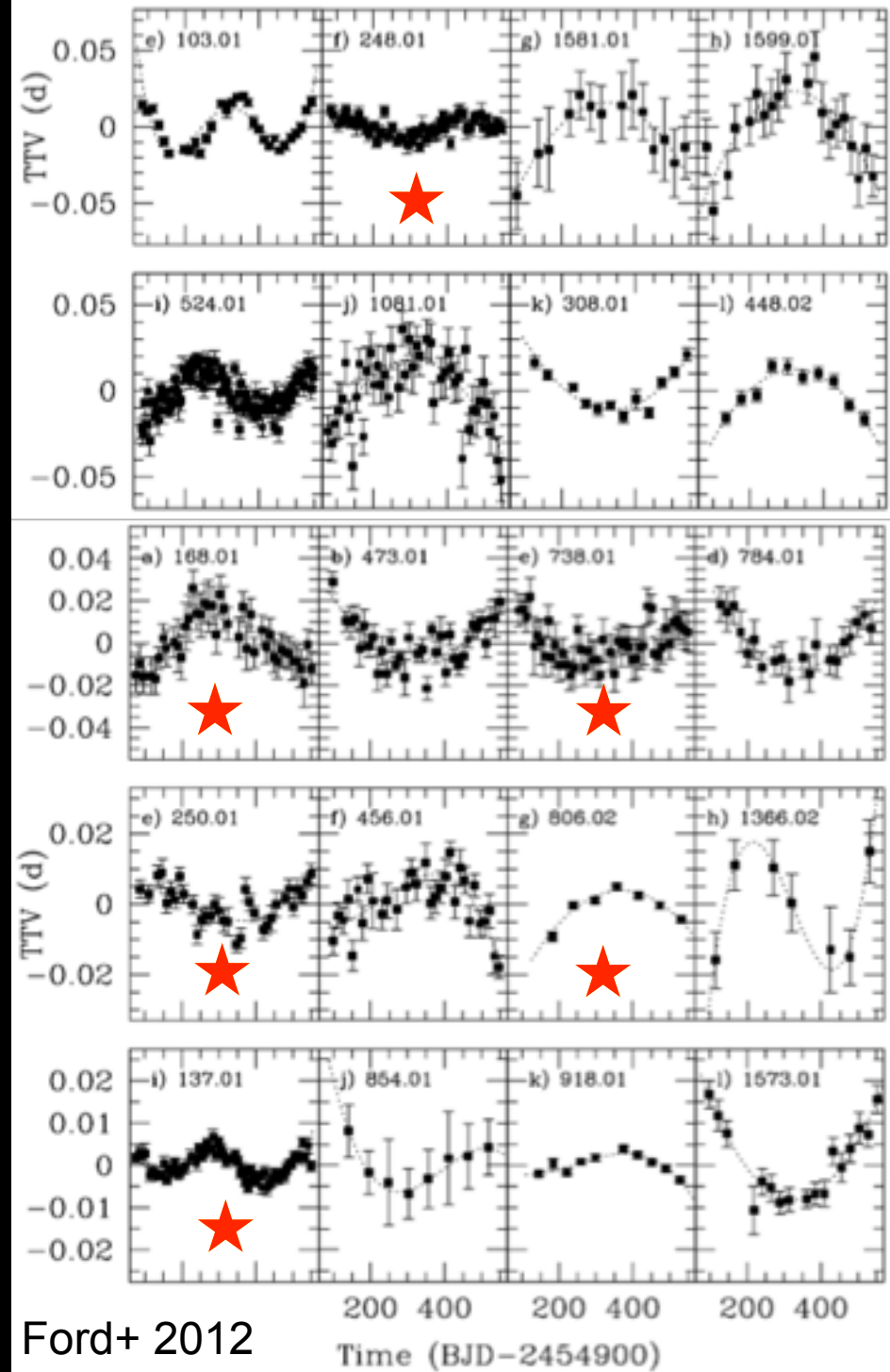
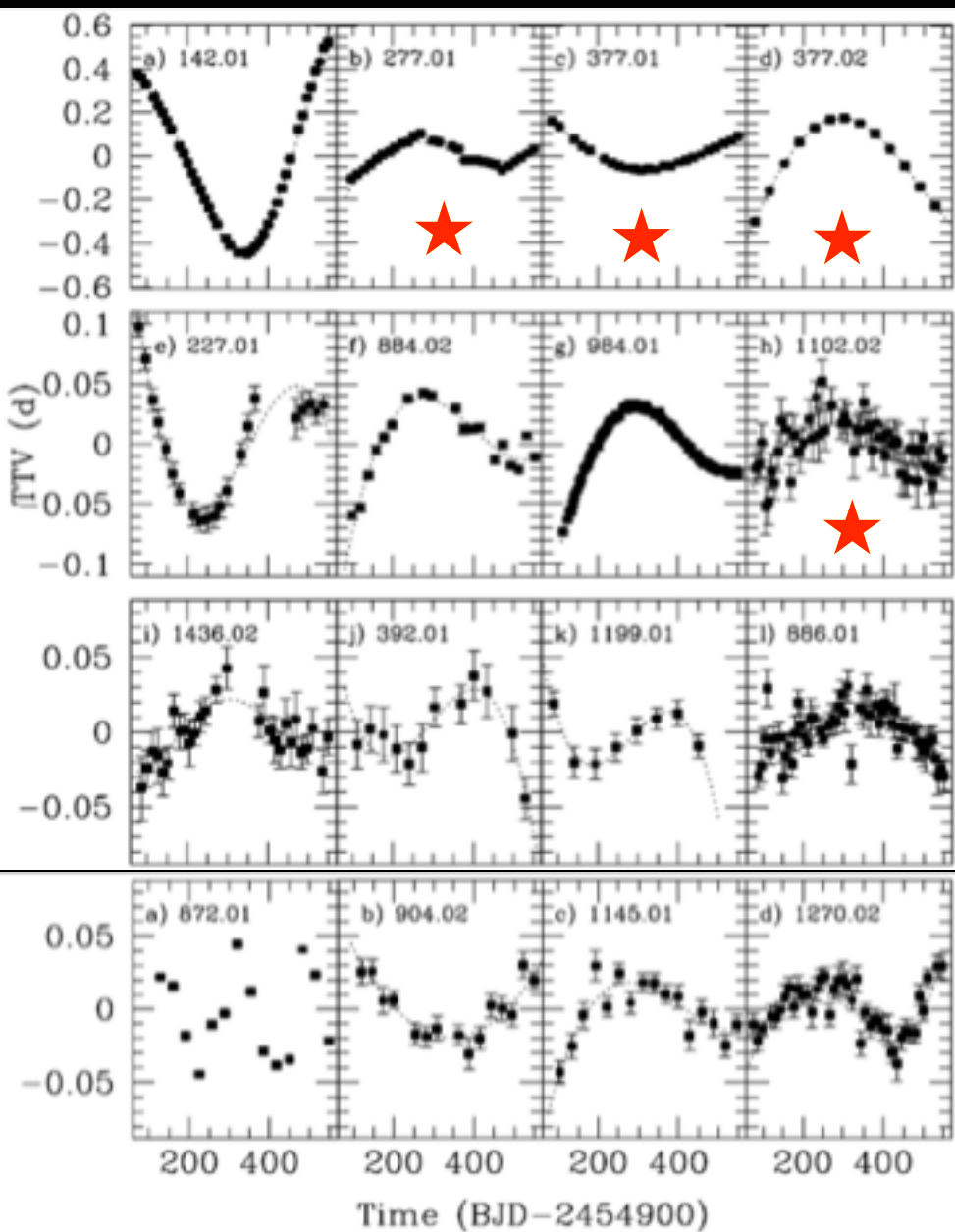
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One complex inverse problem!

(Observables, Desired Distributions, Both)

Long-Term TTVs

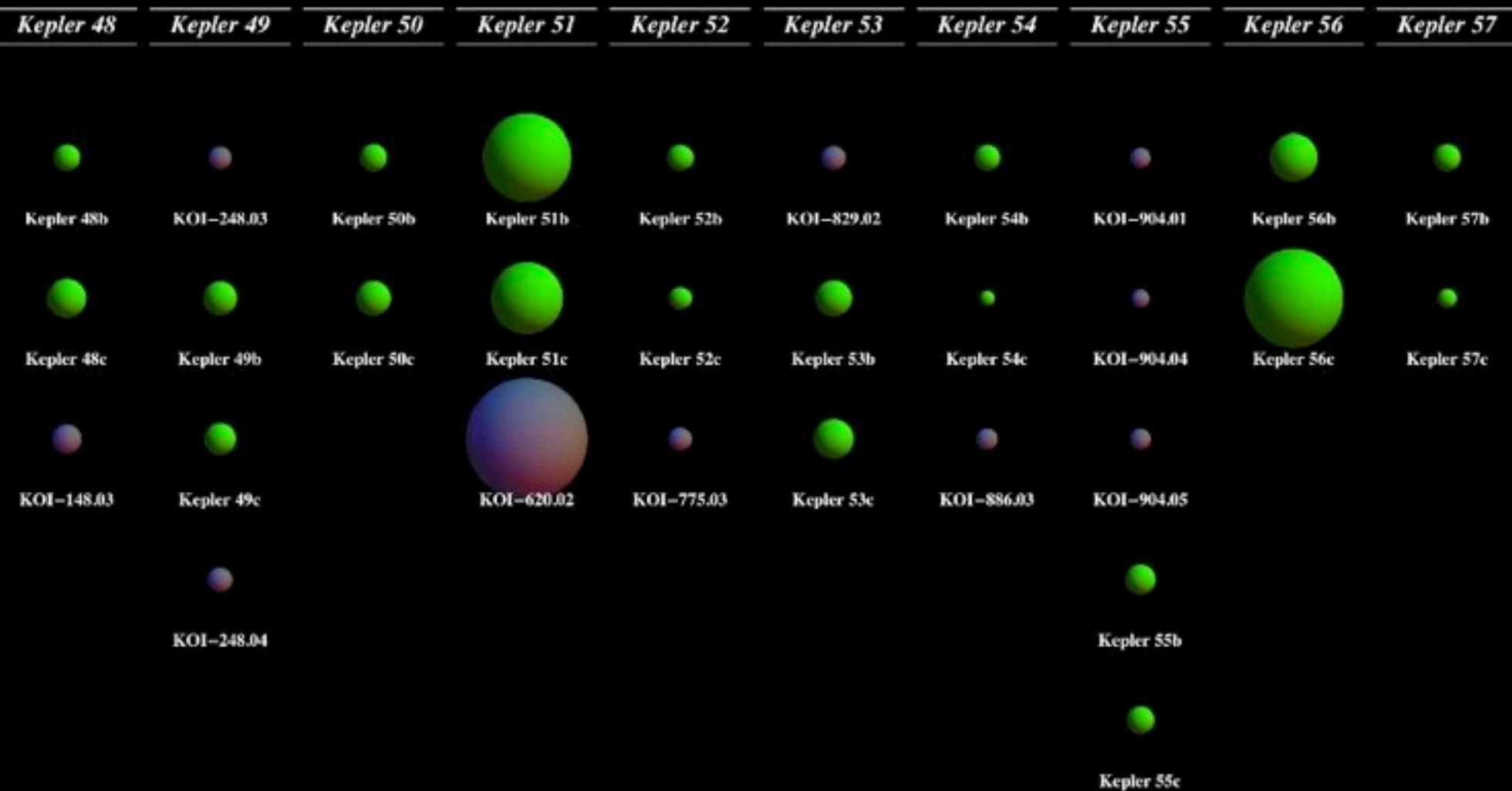


Ford+ 2012

Kepler's Multiple Planet Systems

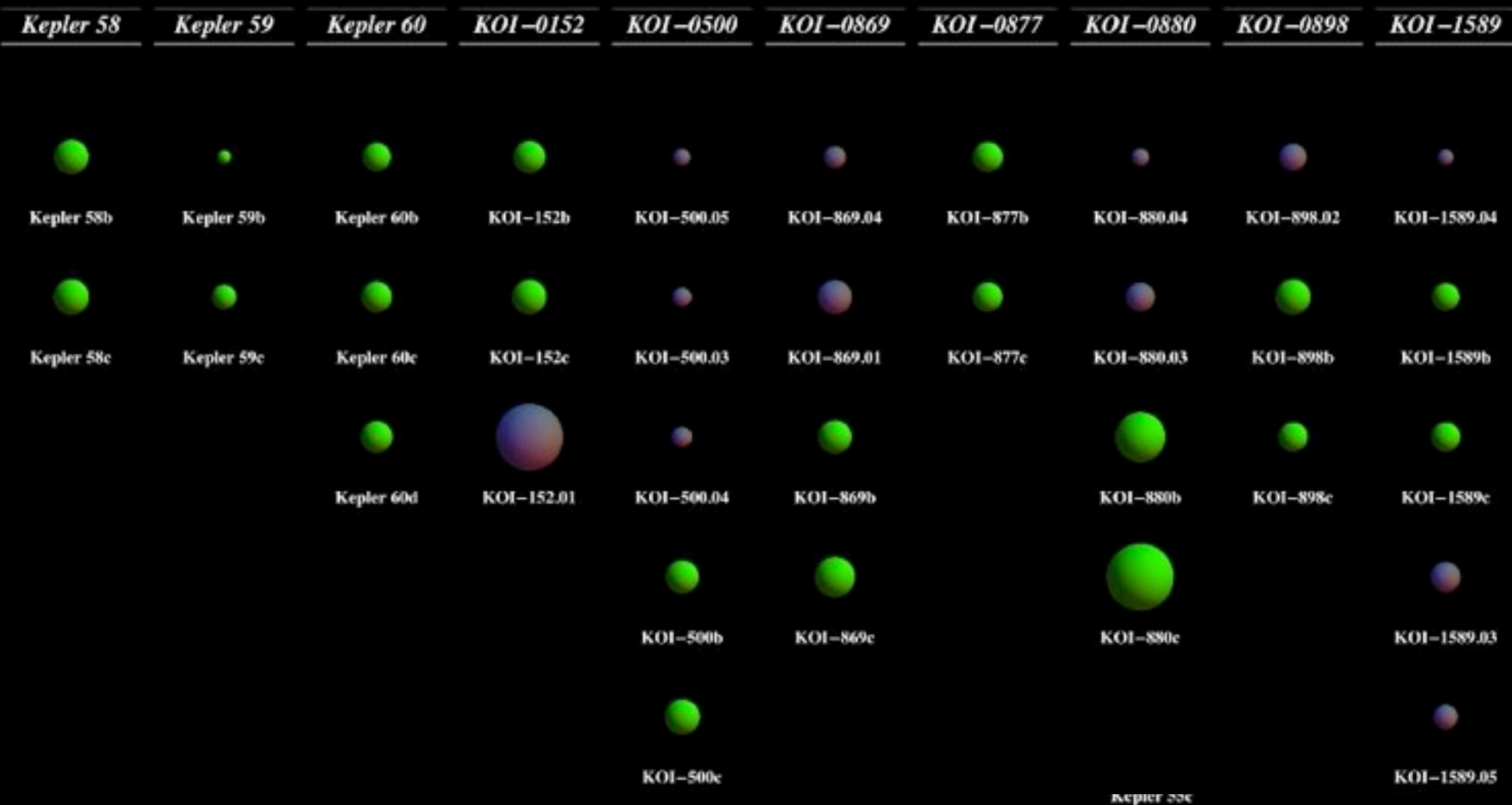
Holman+ 2011; Batalha+ 2010; Torres+ 2010; Lissauer+ 2011; Cochran+ 2011; Ford+ 2012; Steffen+ 2012ab;
Fabrycky+ 2012; Fressin+ 2012; Muirhead+ 2012; Nesvorny+ 2012; Xie 2012; Orosz+ 2012

Kepler's Multiple Planet Systems



Holman+ 2011; Batalha+ 2010; Torres+ 2010; Lissauer+ 2011; Cochran+ 2011; Ford+ 2012; Steffen+ 2012ab; Fabrycky+ 2012; Fressin+ 2012; Muirhead+ 2012; Nesvorny+ 2012; Xie 2012; Orosz+ 2012

Kepler's Multiple Planet Systems



Holman+ 2011; Batalha+ 2010; Torres+ 2010; Lissauer+ 2011; Cochran+ 2011; Ford+ 2012; Steffen+ 2012ab; Fabrycky+ 2012; Fressin+ 2012; Muirhead+ 2012; Nesvorny+ 2012; Xie 2012; Orosz+ 2012

Kepler's Multiple Planet Systems

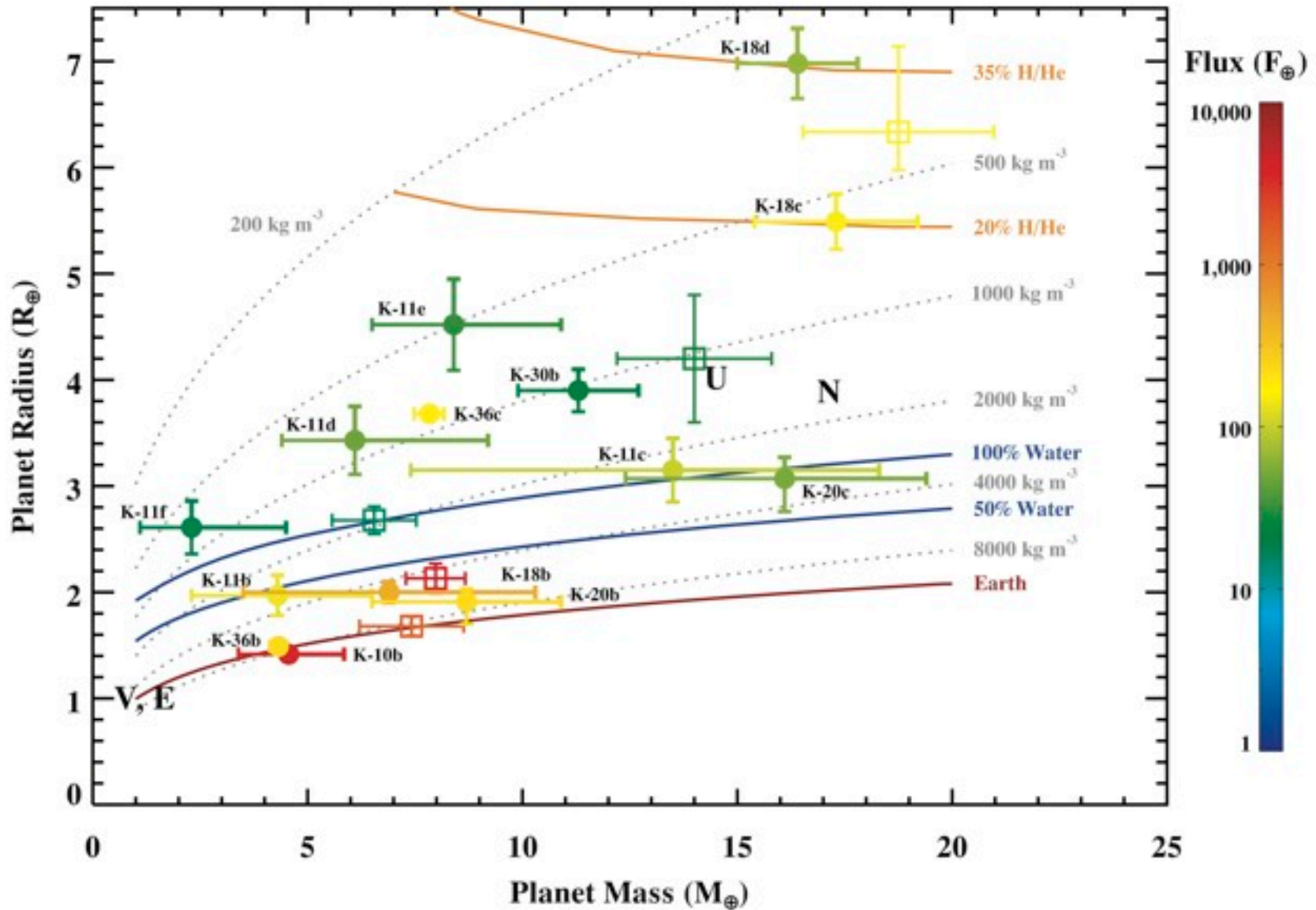


Holman+ 2011; Batalha+ 2010; Torres+ 2010; Lissauer+ 2011; Cochran+ 2011; Ford+ 2012; Steffen+ 2012ab; Fabrycky+ 2012; Fressin+ 2012; Muirhead+ 2012; Nesvorny+ 2012; Xie 2012; Orosz+ 2012

How Common are TTVs?

- First 16 months of Kepler data (Ford+ 2012):
 - 39 – 175 TTV candidates
 - 8% – 18% of suitable KOIs show TTVs
 - More for multis & long period planets
- Planets Confirmed by TTVs: 73 of ~105
- Sensitivity to long-term TTVs grows $\sim t^{5/2}$
 - Many more KOIs w/ TTVs in extended mission

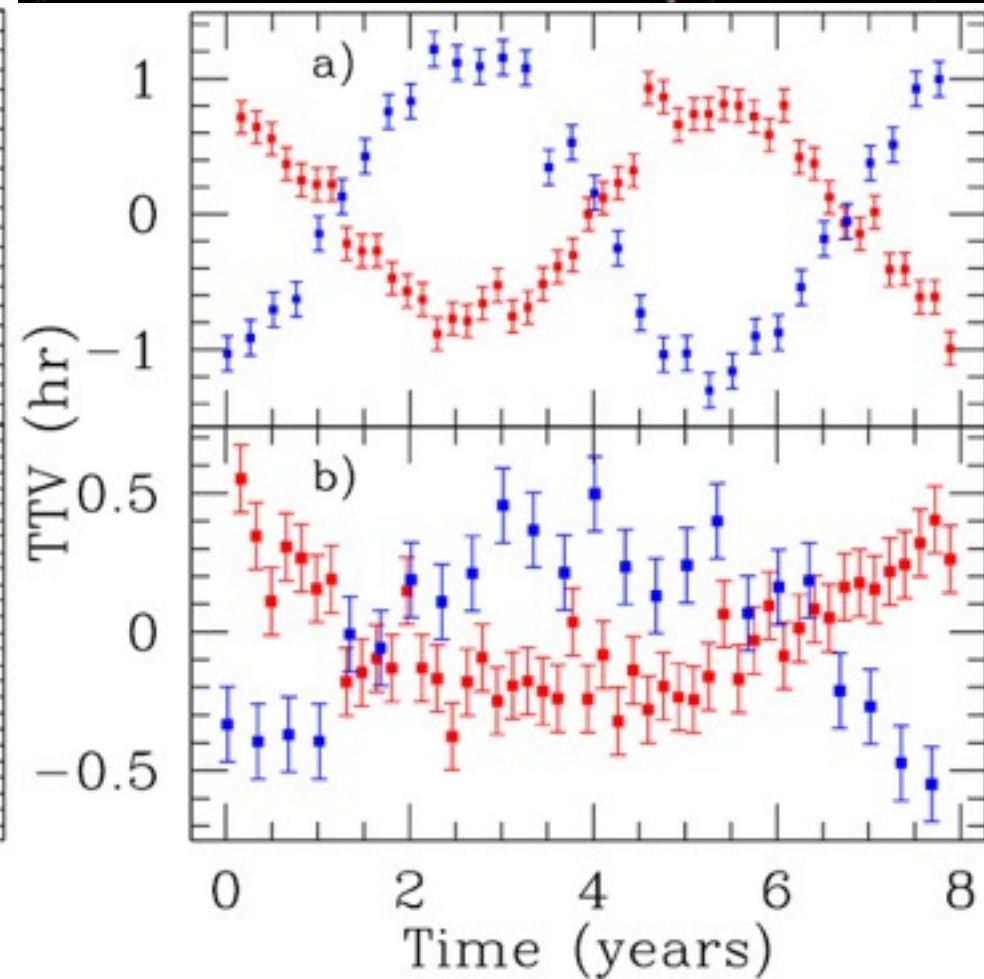
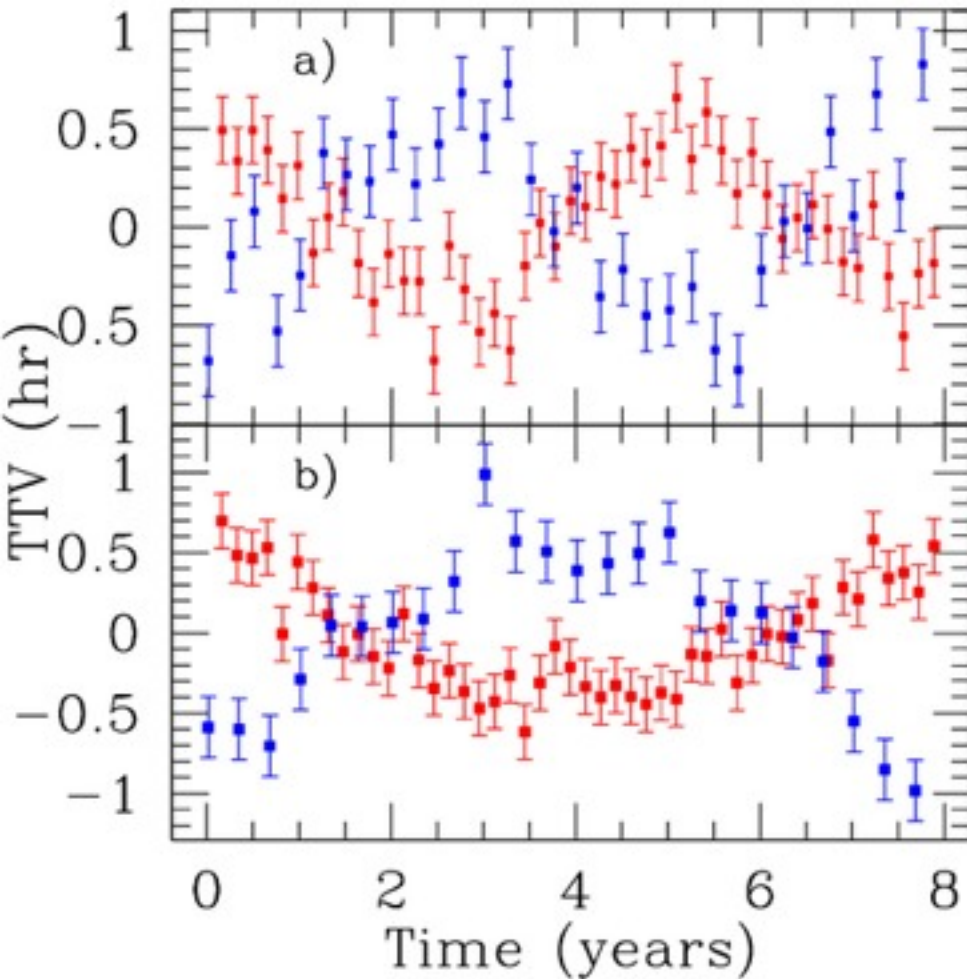
Super-Earths or Mini-Neptunes?



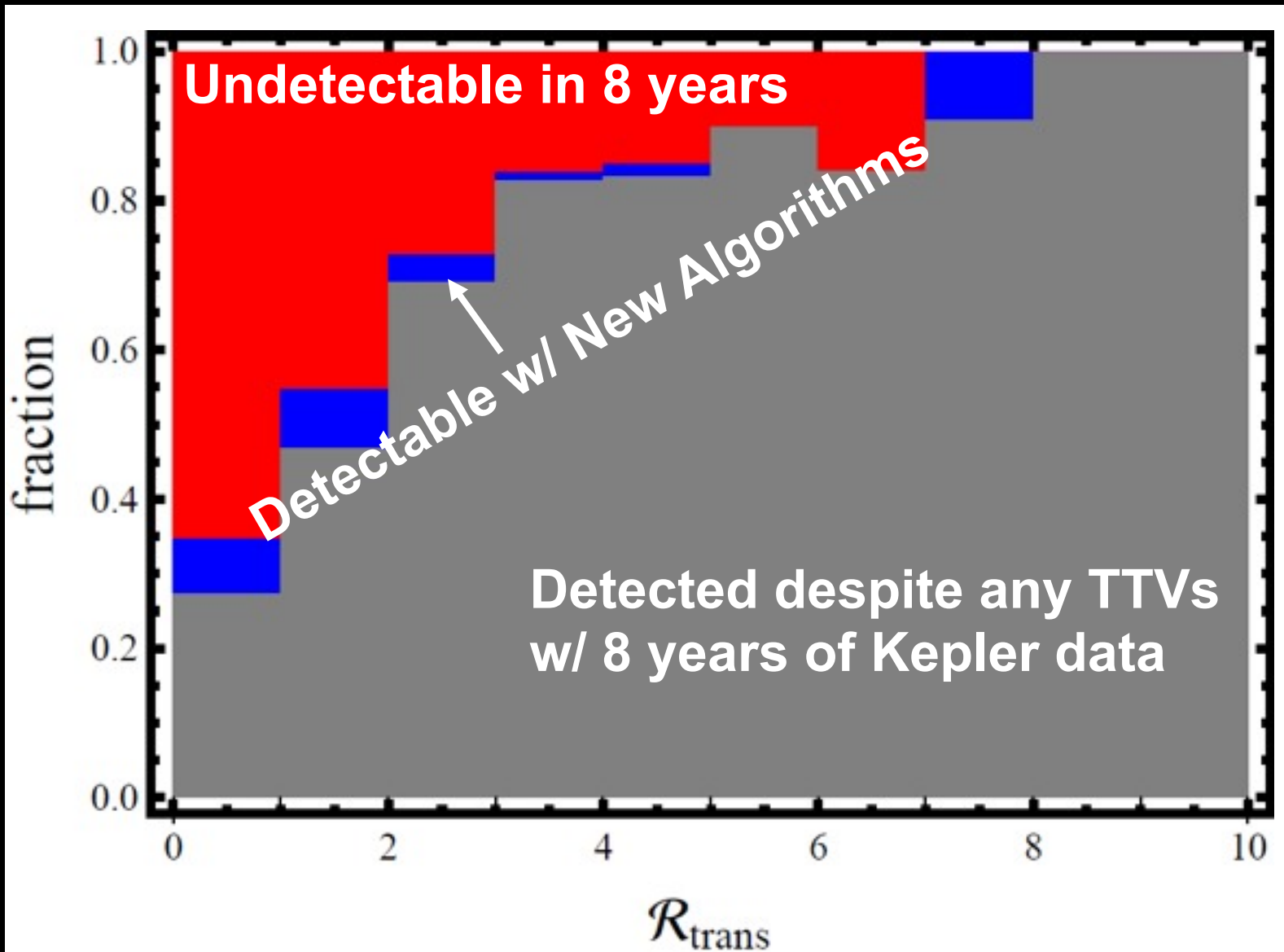
Future Prospects for Measuring Masses via TTVs

1 Earth-mass, 3:2 MMR, $K_p=13$

2 Earth-mass, 3:2 MMR, $K_p=13$



Detecting Small Planets w/ Large TTVs

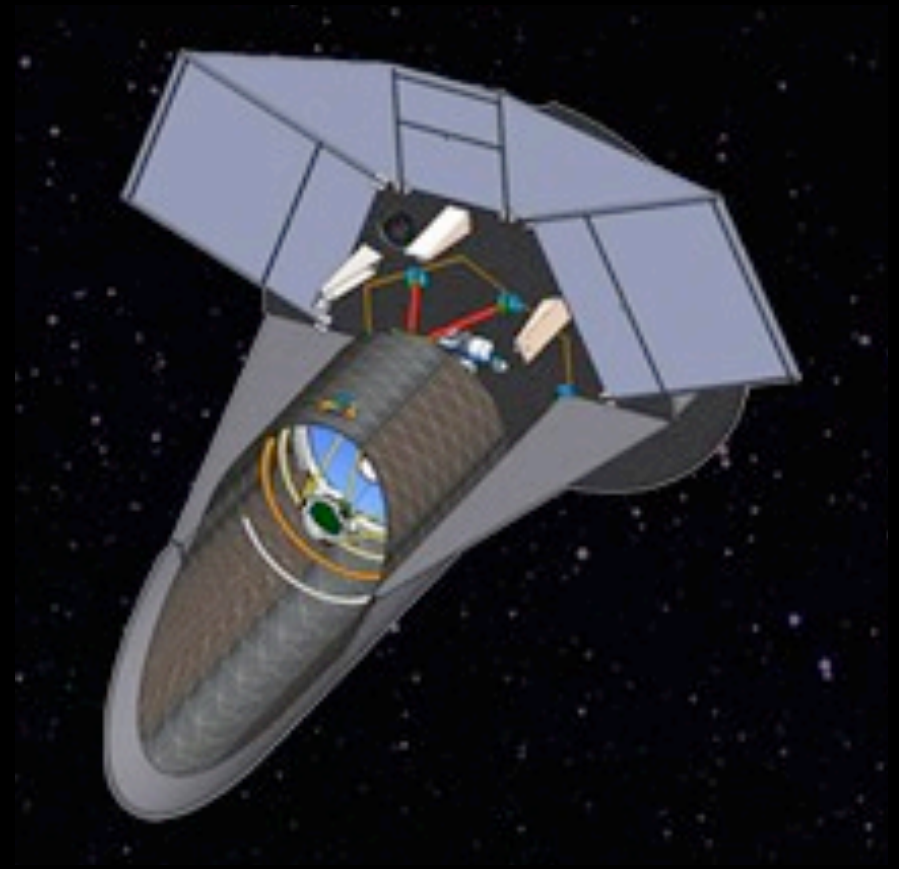
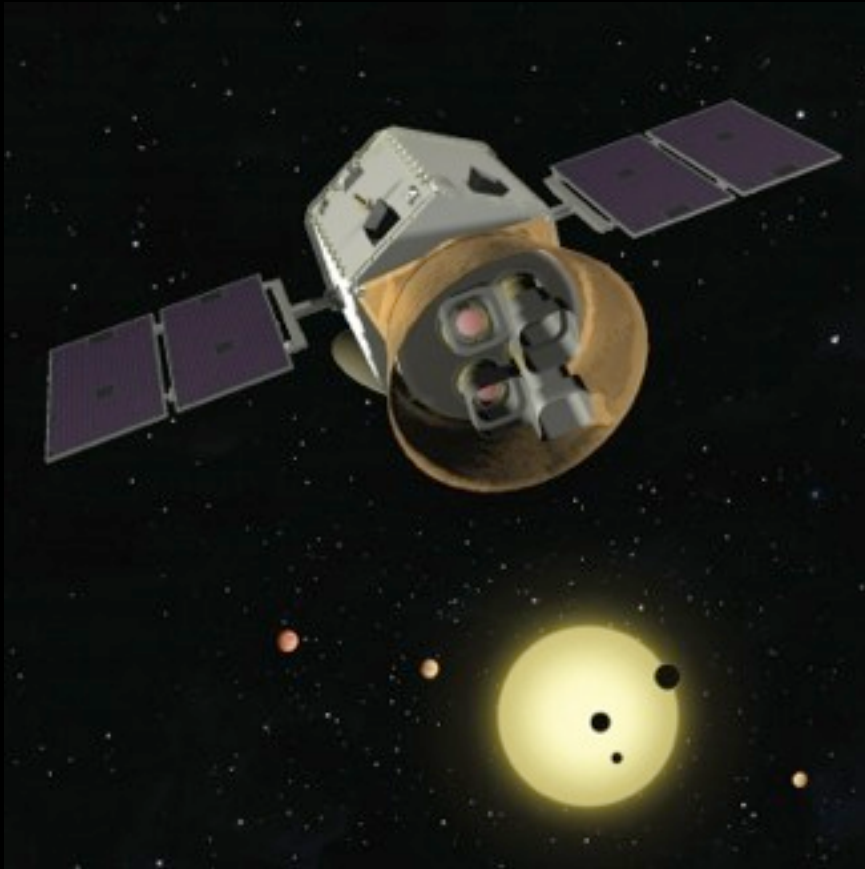


Testing Planet Formation Theory with Kepler

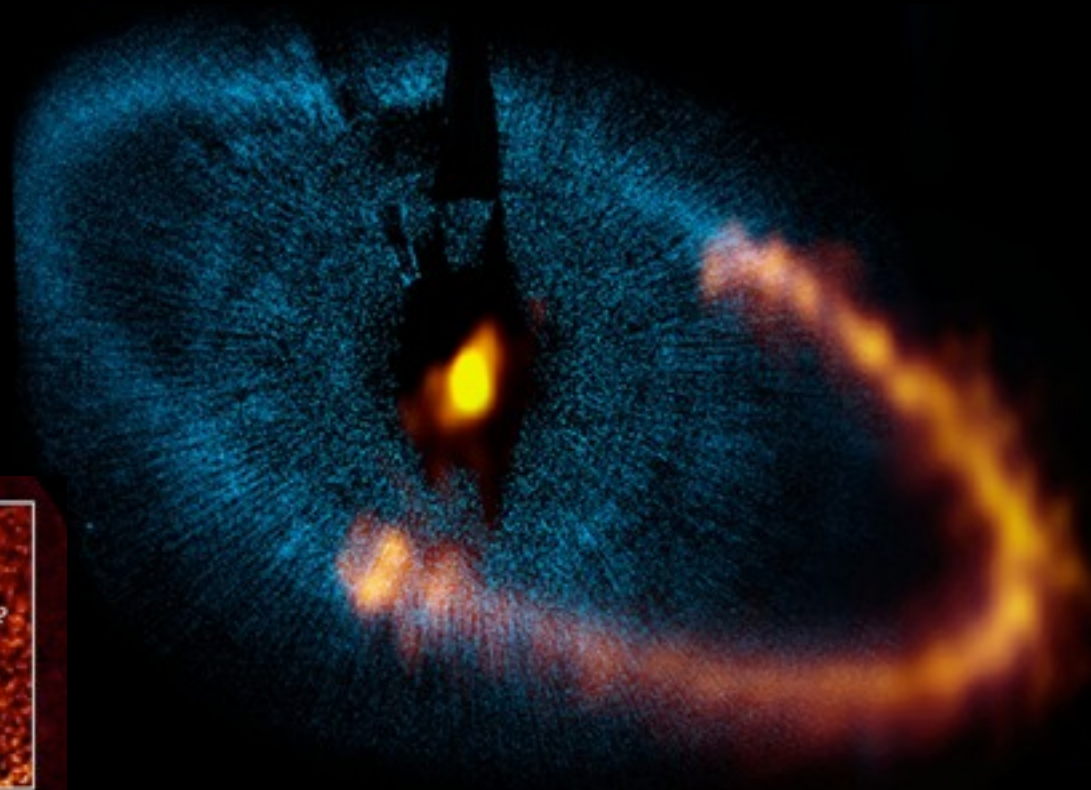
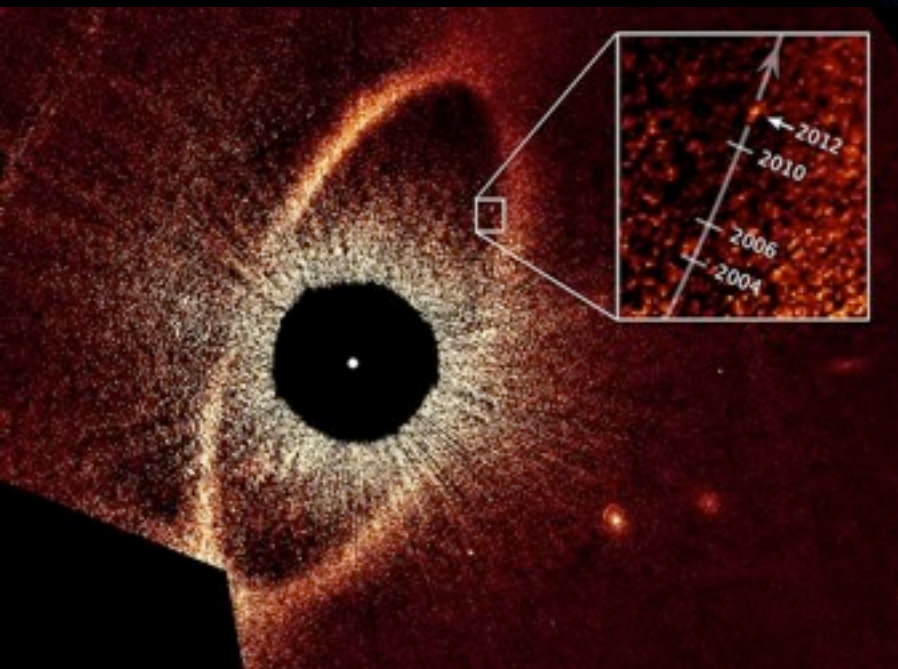
Must combine many elements simultaneously:

- Detection efficiency/completeness
- Planetary systems
(not just superposition of individual planets)
- Variety of observational constraints
(e.g., RV, TTV, spectra, imaging, seismology)
- Observational uncertainties
- How planets were chosen for follow-up observations or more detailed analyses

Future Space Missions

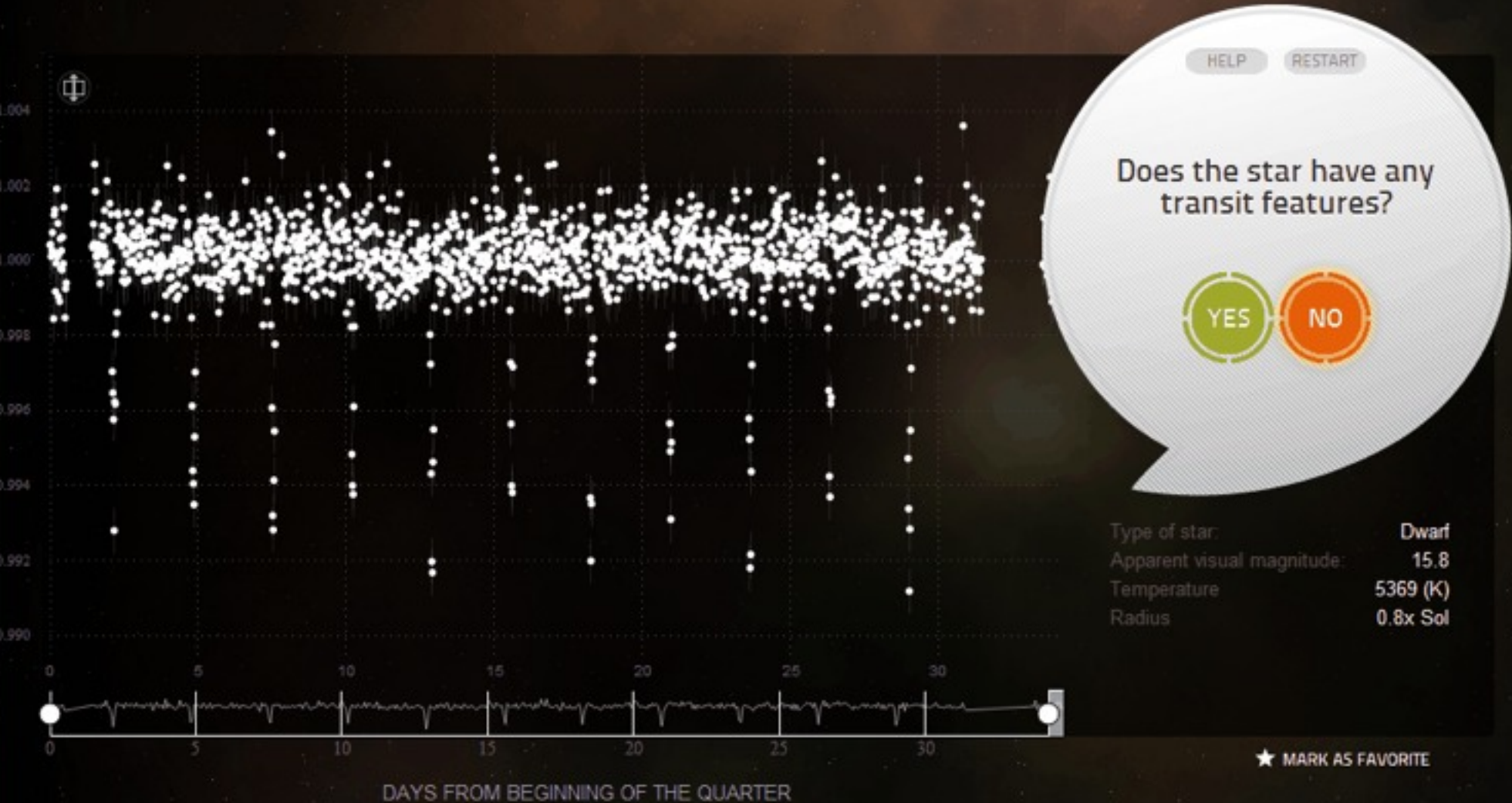


Direct Imaging & ALMA

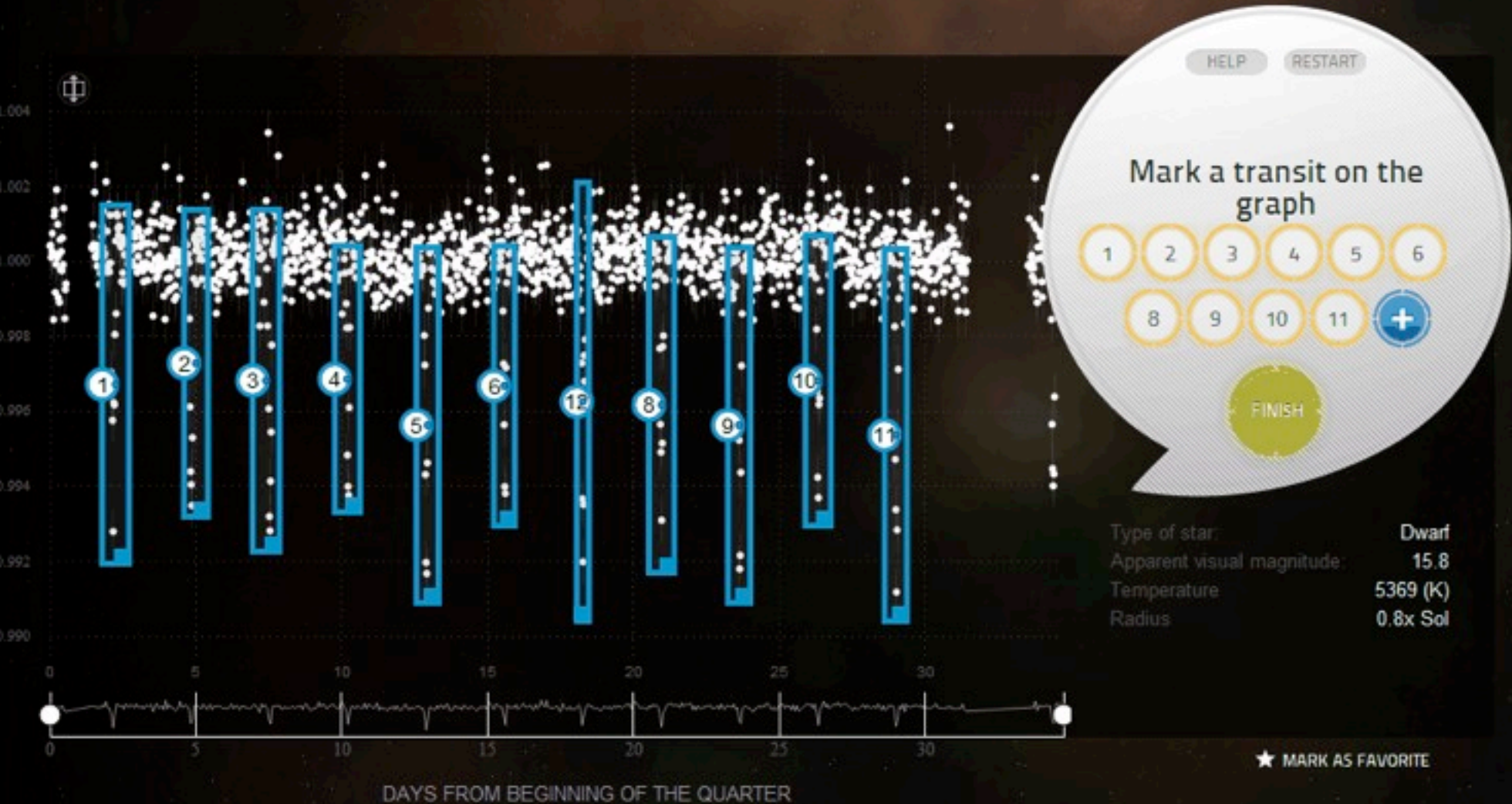


Boley et al. 2012
ALMA: ESO/NAOJ/NRAO + HST: NASA/ESA

Invite Students to Join the Hunt!



Invite Students to Join the Hunt!



Questions

The logo for the Kepler mission is centered on the page. It features a dark silhouette of a bird, resembling a starling, with several bright yellow stars scattered across its body and wings. Below the bird, the word "Kepler" is written in a large, blue, 3D-style font. A yellow orbital path is drawn around the letters, with a small blue and white planet (representing Earth) positioned at one point on the orbit. The background is a dark space filled with numerous small white stars, and the top edge of the Earth's horizon is visible at the bottom of the frame.

Kepler

Movie of Collapse & Fragmentation

