

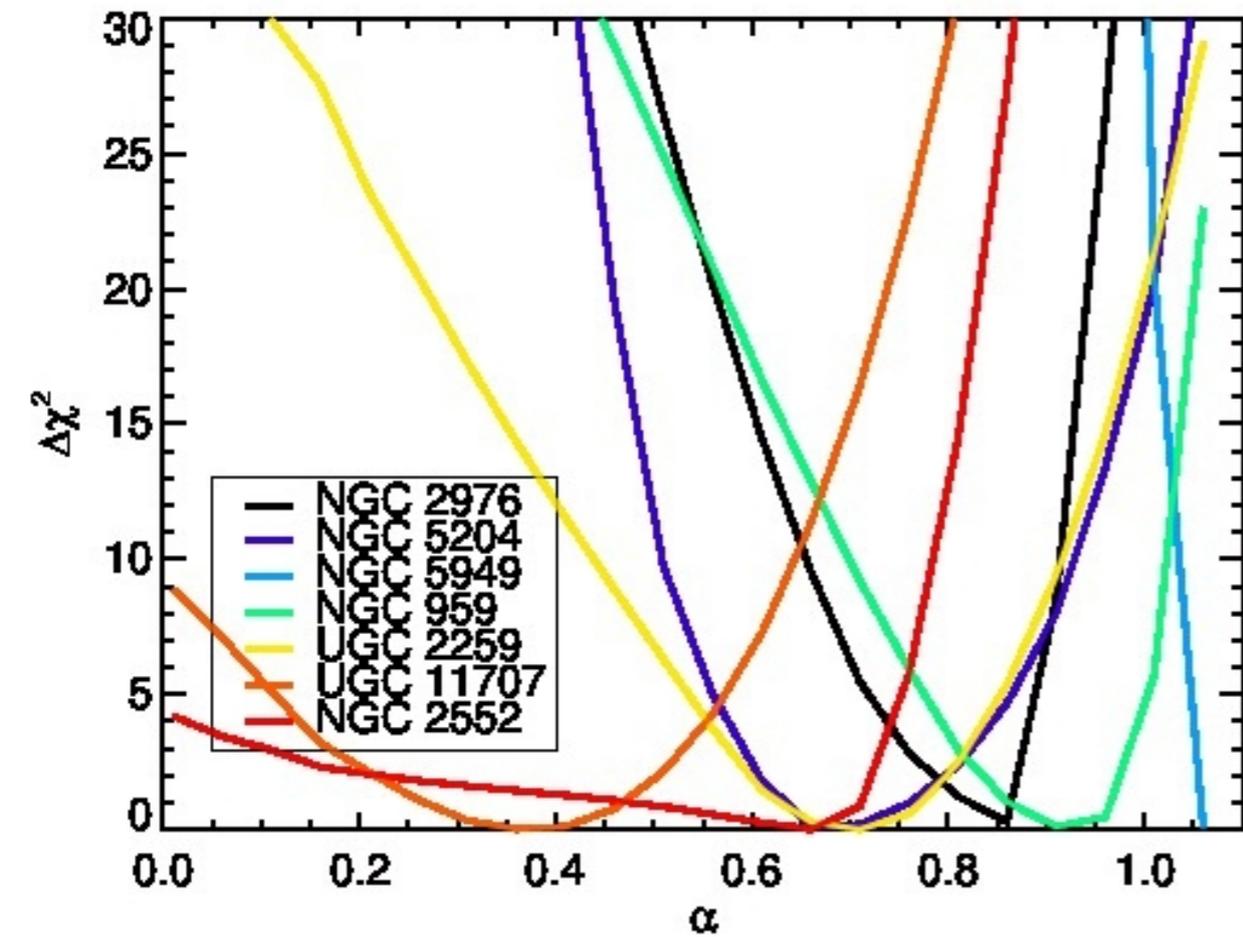
The MEDIDO project: surprising result for Leo I



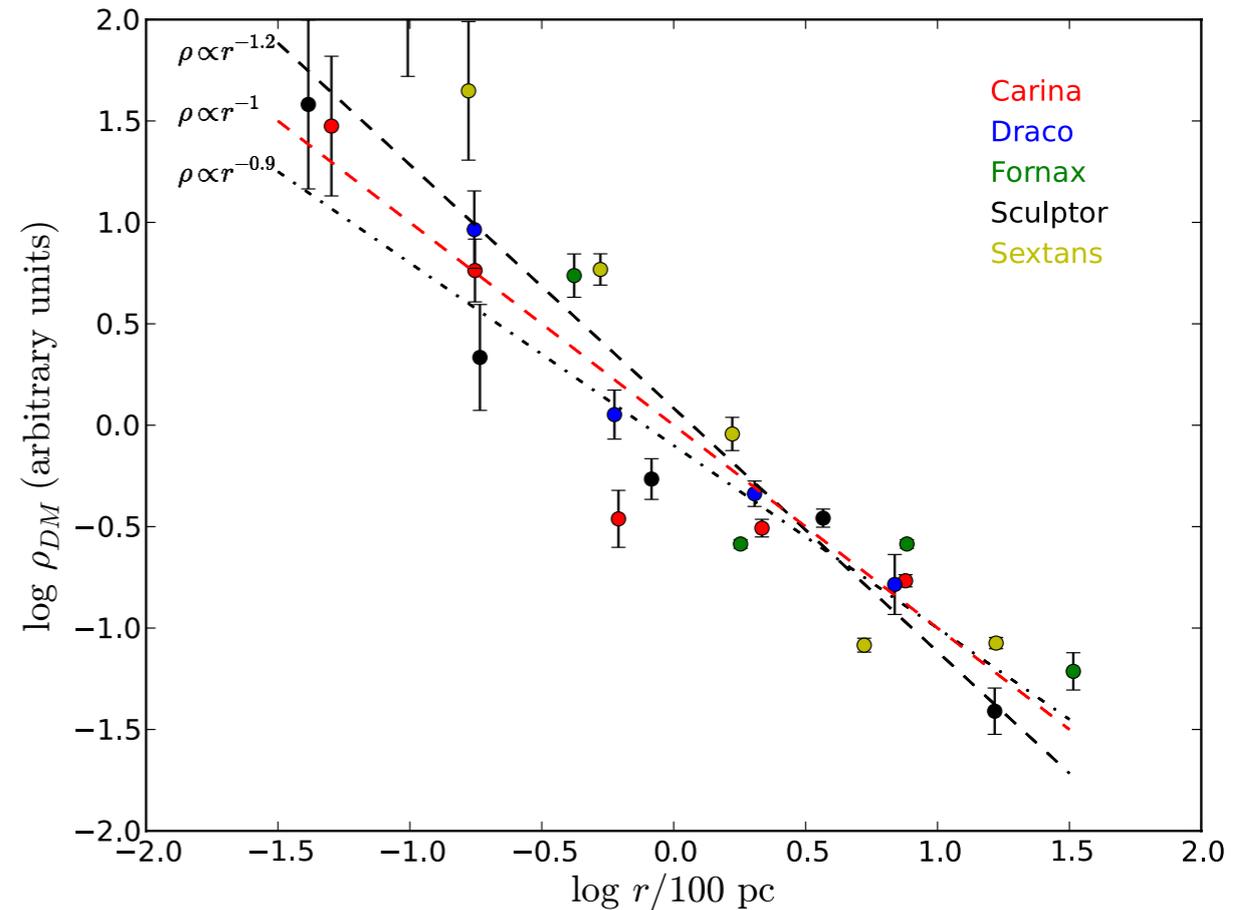
Eva Noyola
McDonald Observatory

Majo Bustamante-Rosell, Karl Gebhardt (UT, Austin)

Shape of dark matter haloes in dwarf galaxies



Adams et al, 2013

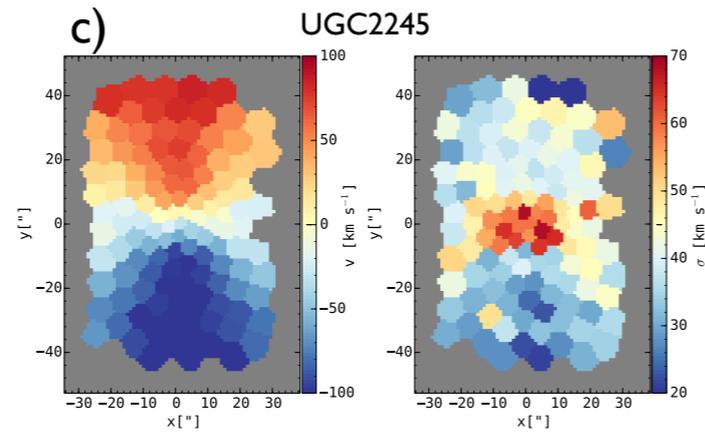
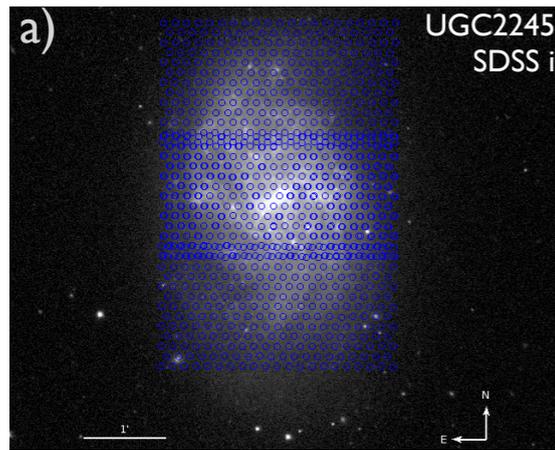


Jardel & Gebhardt, 2013

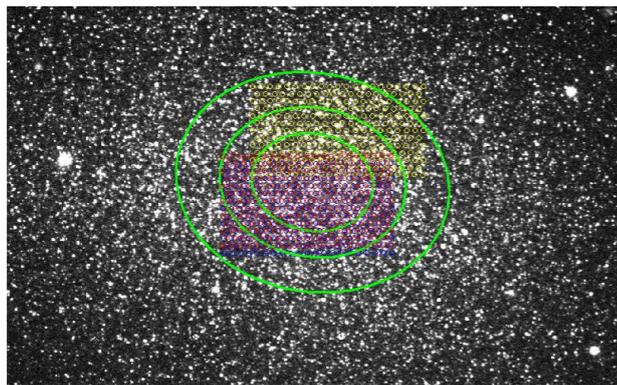
- Core/cusp problem: measure the central shape of DM halos
- Missing Satellite problem: measure accurate velocities for dwarf galaxies and test DM content for globular clusters

MEDIDO (MEasuring Dynamics In Dwarf Objects)

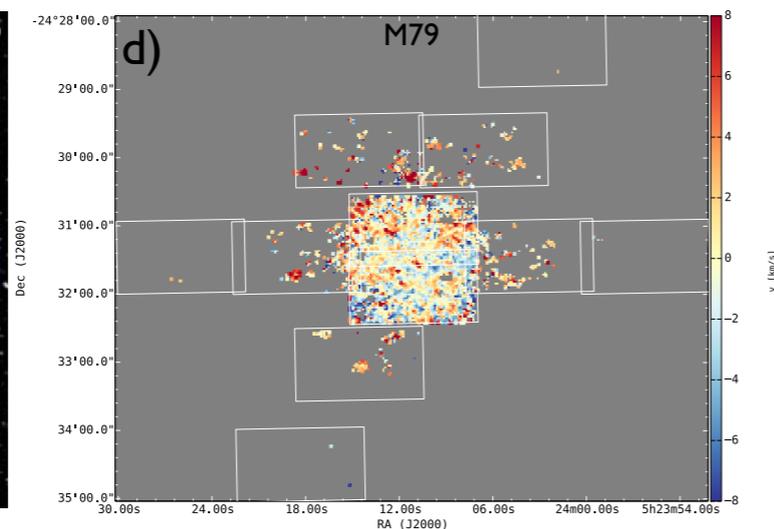
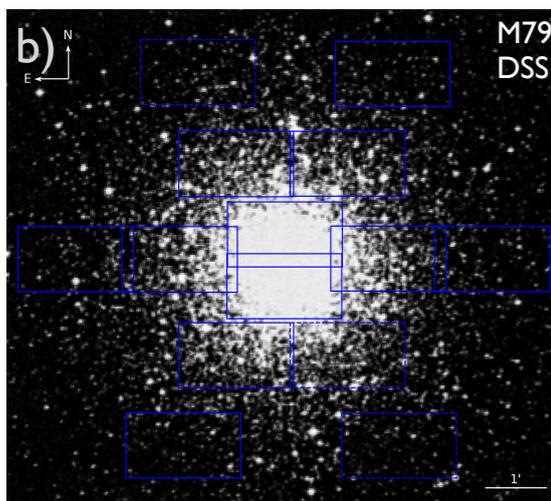
Sample



Unresolved dwarf galaxies

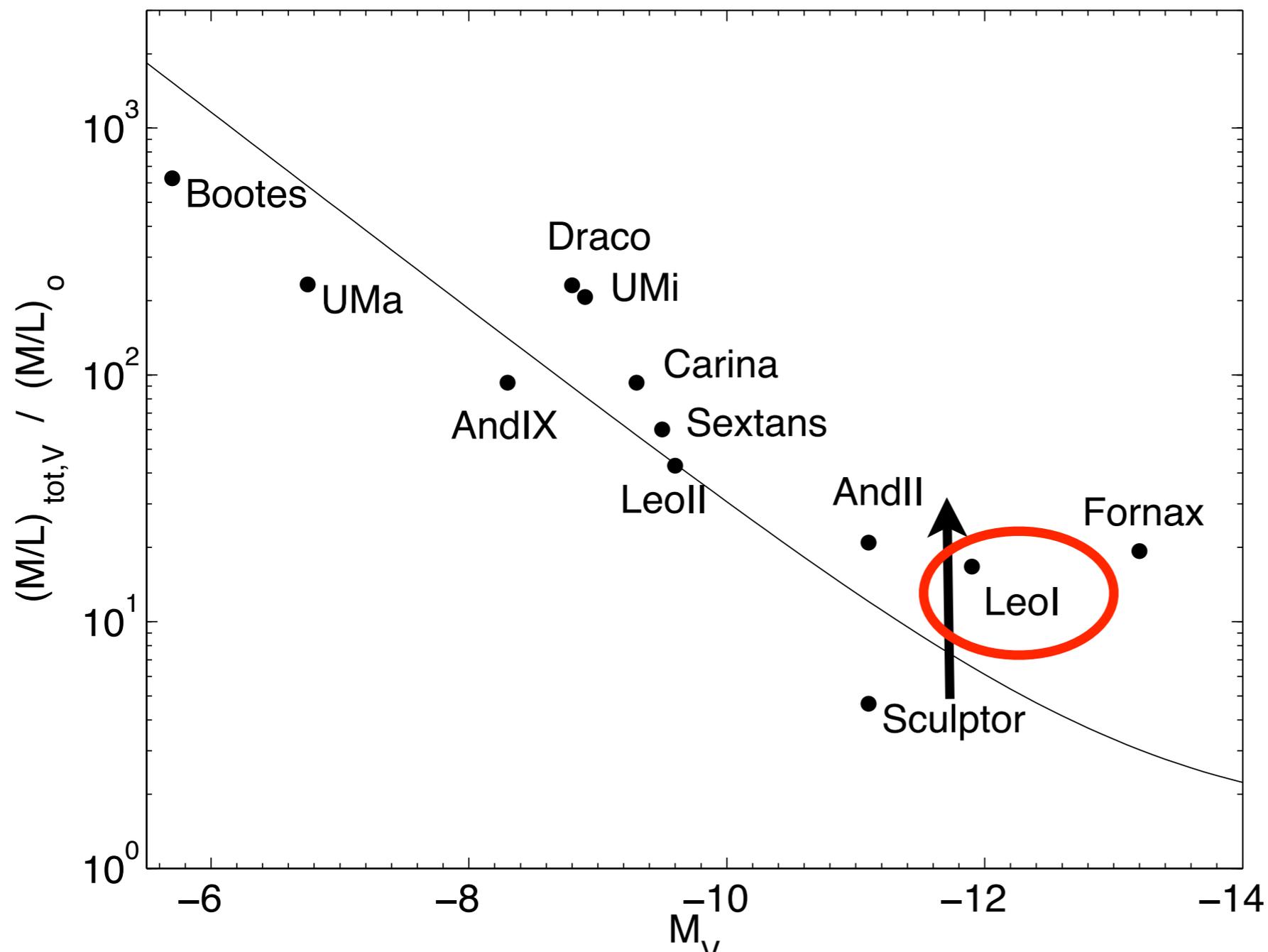


Resolved dwarf galaxies



Galactic globular clusters

Leo I among MW dwarf spheroidals



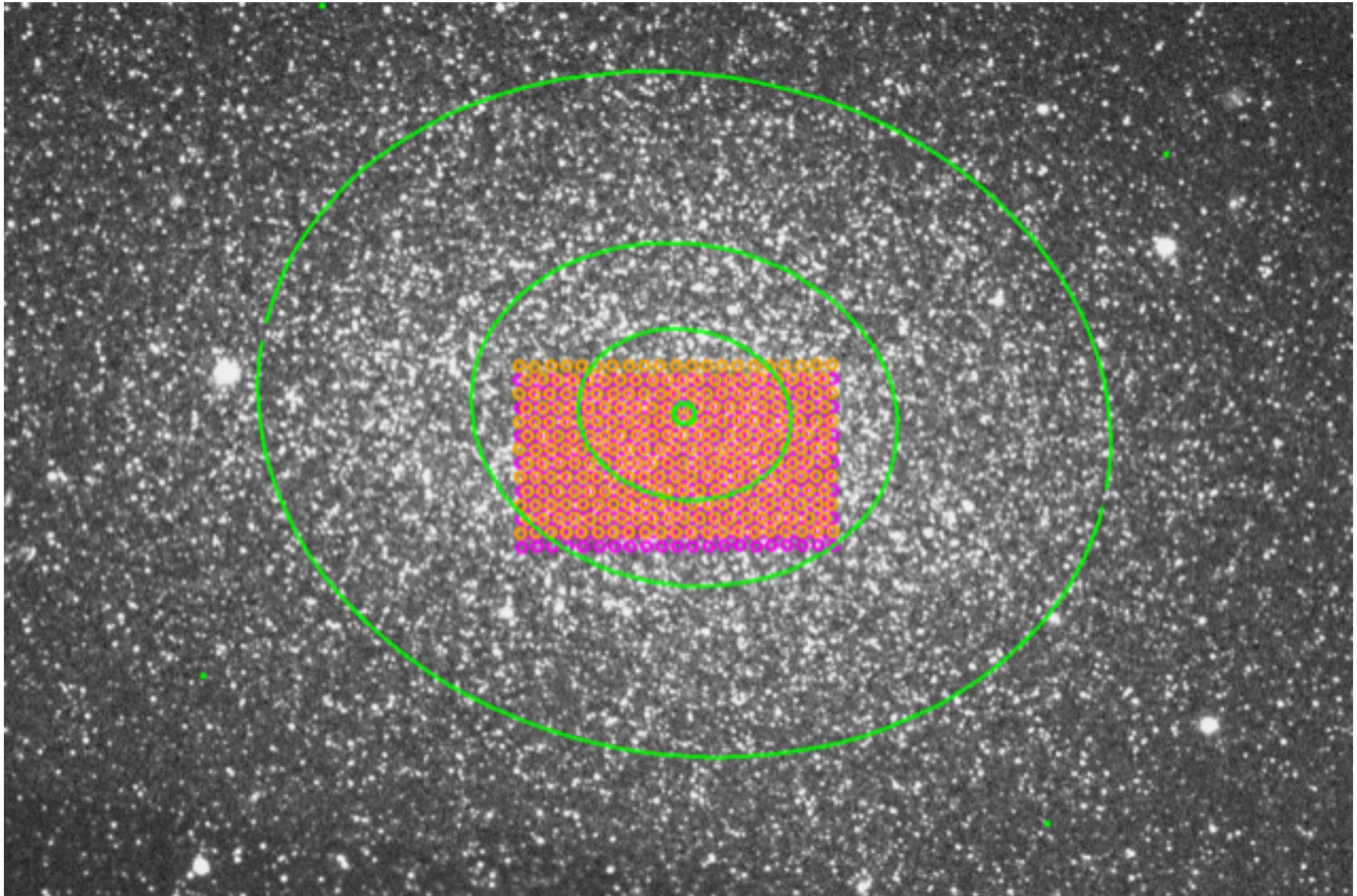
Mateo et al. 1998

Wilkinson et al 2006

Gilmore et al. 2007

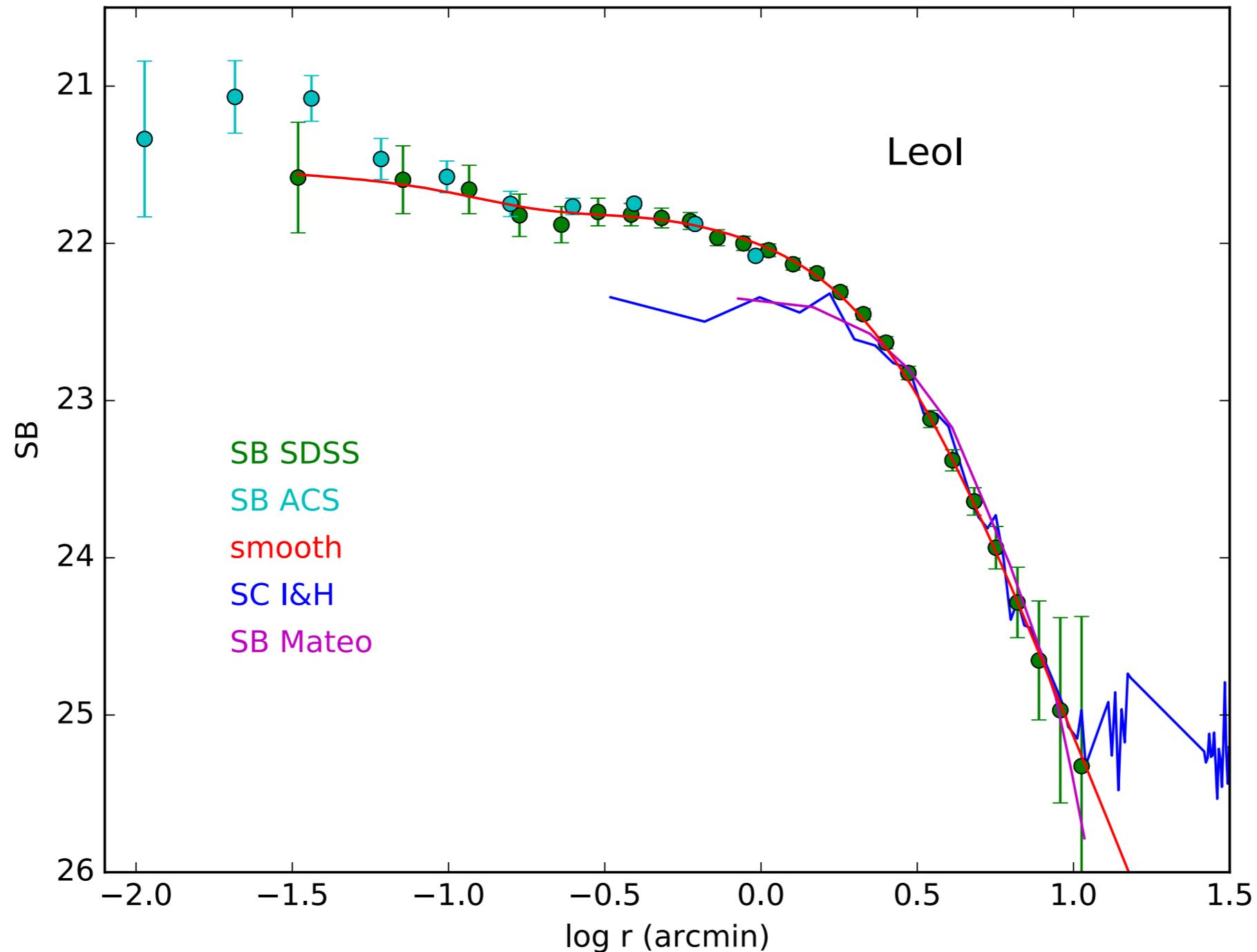
- Leo I is almost the richest MW dwarf galaxy, and it has a low measured M/L

VIRUS-W coverage



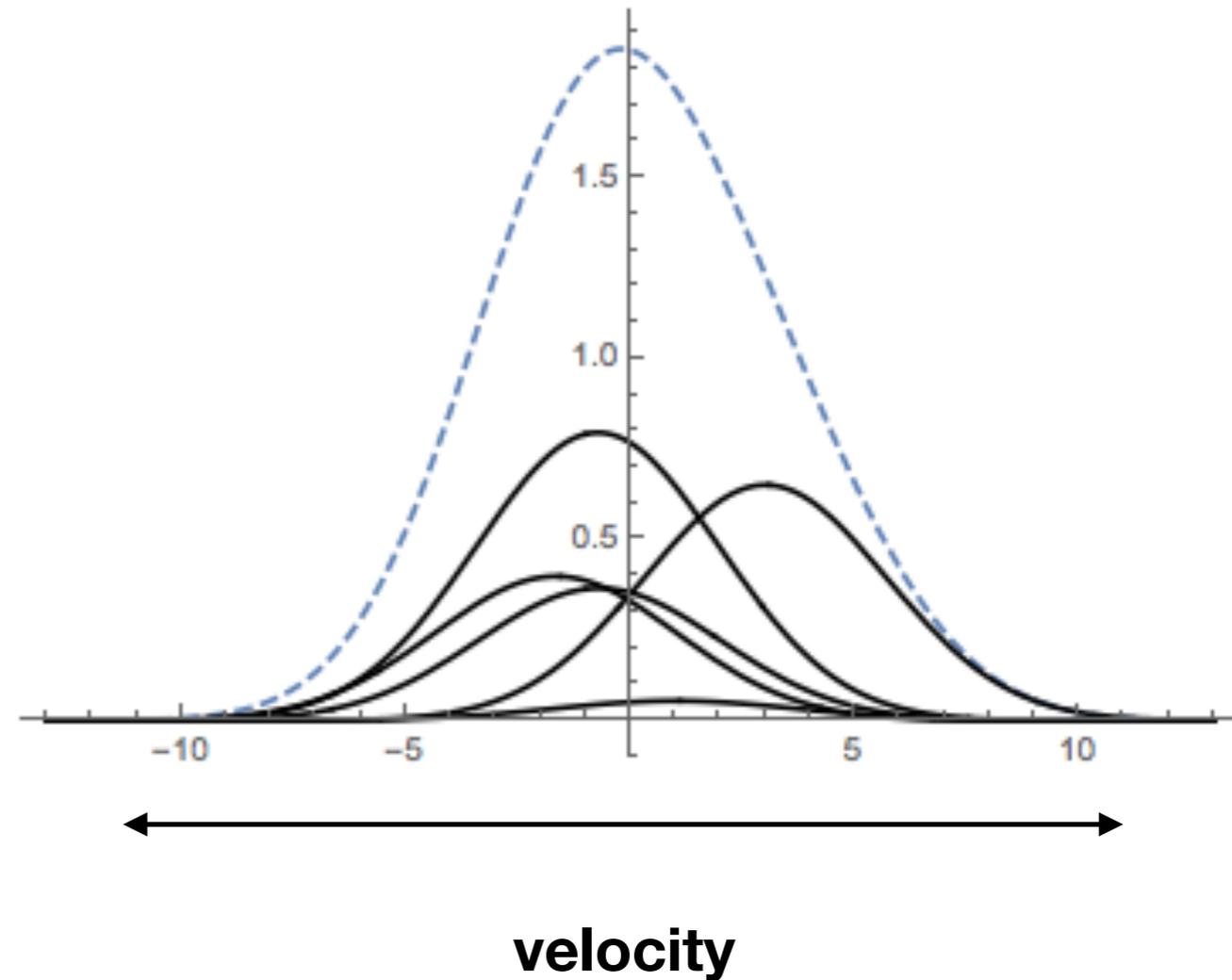
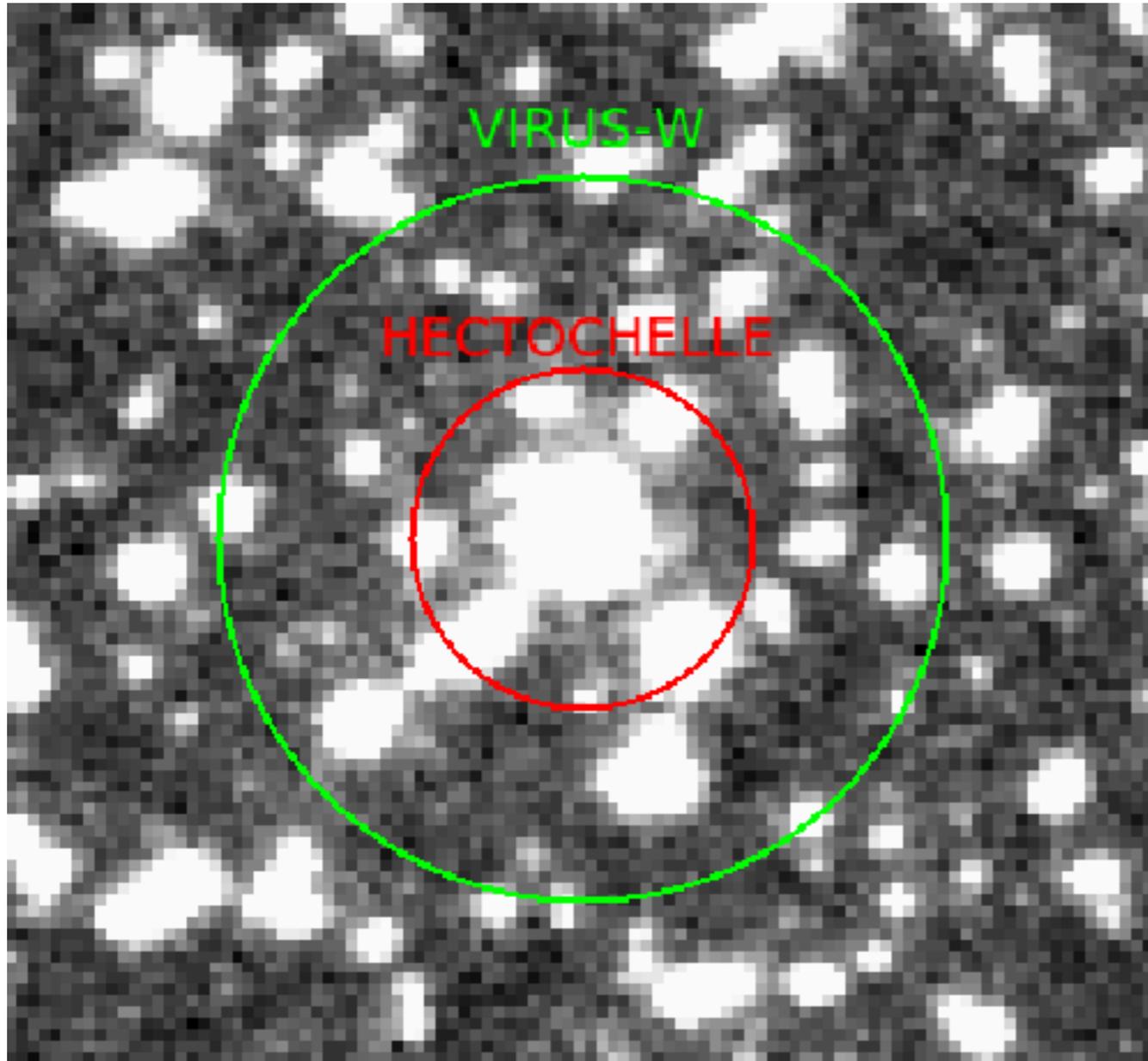
- Previously only a couple dozen stars measured in the central region by Mateo et al. (1998)

Updated density profile



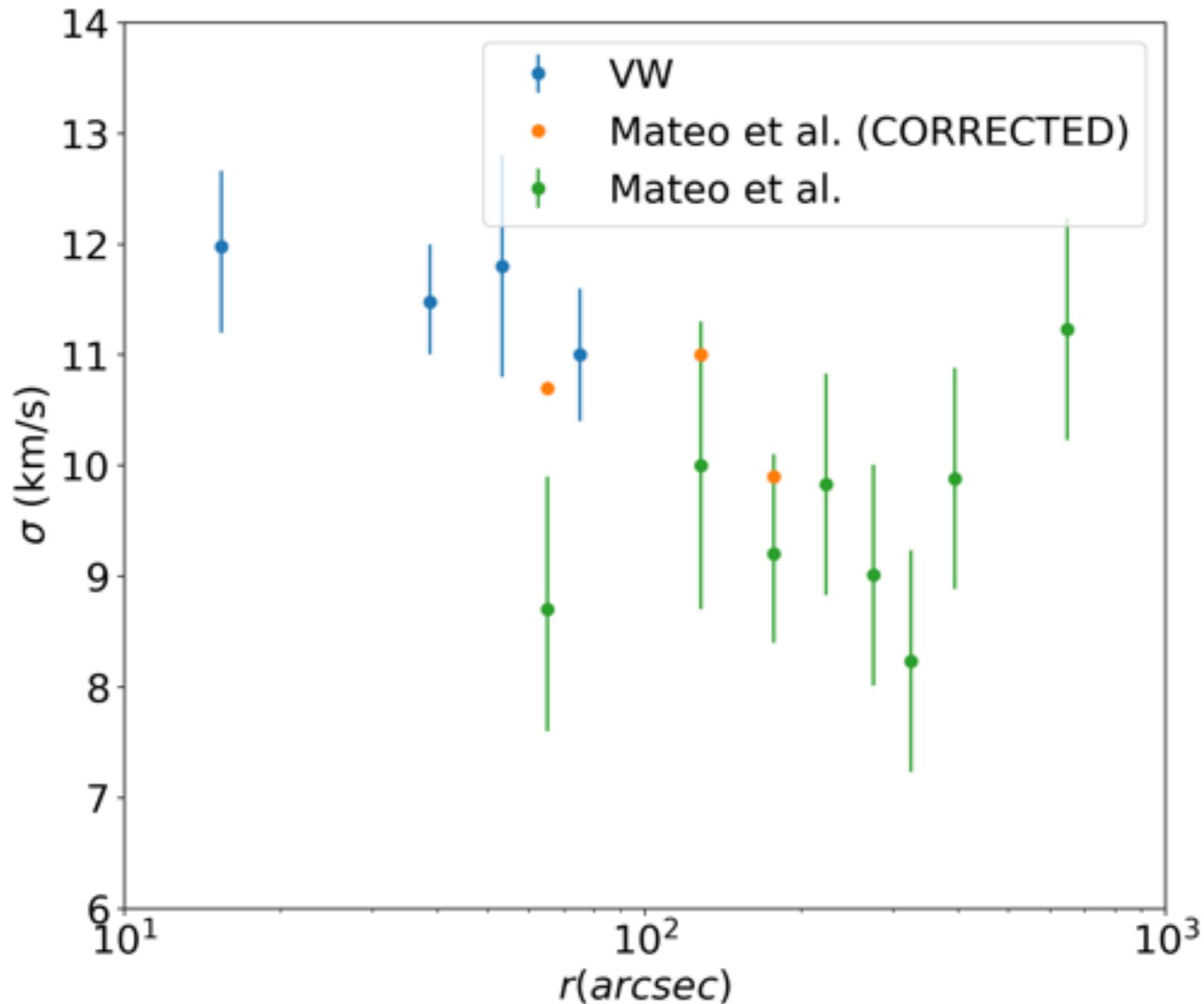
- Central surface brightness profile shows a shallow cusp, both for ground based and HST data

Kinematics with virus-w IFU



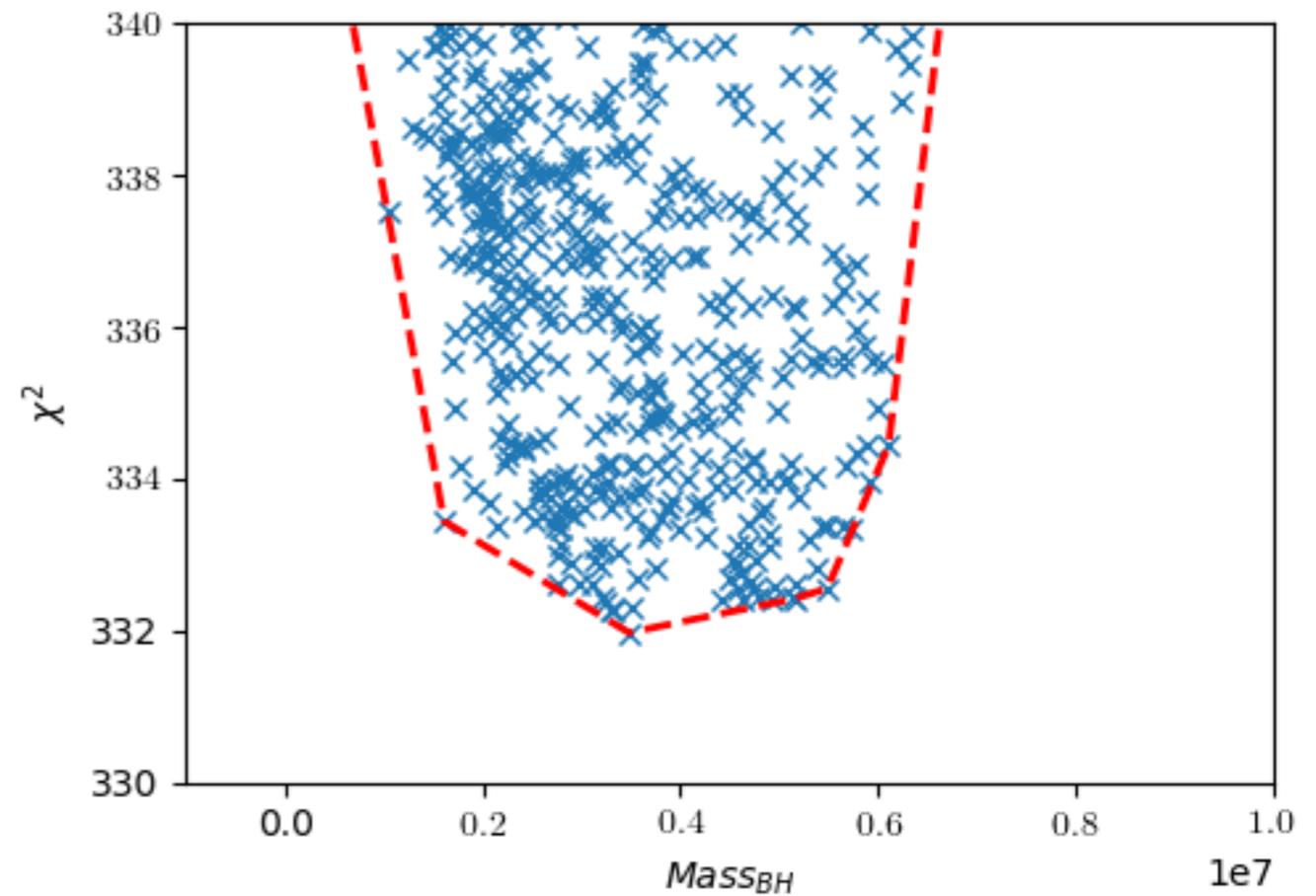
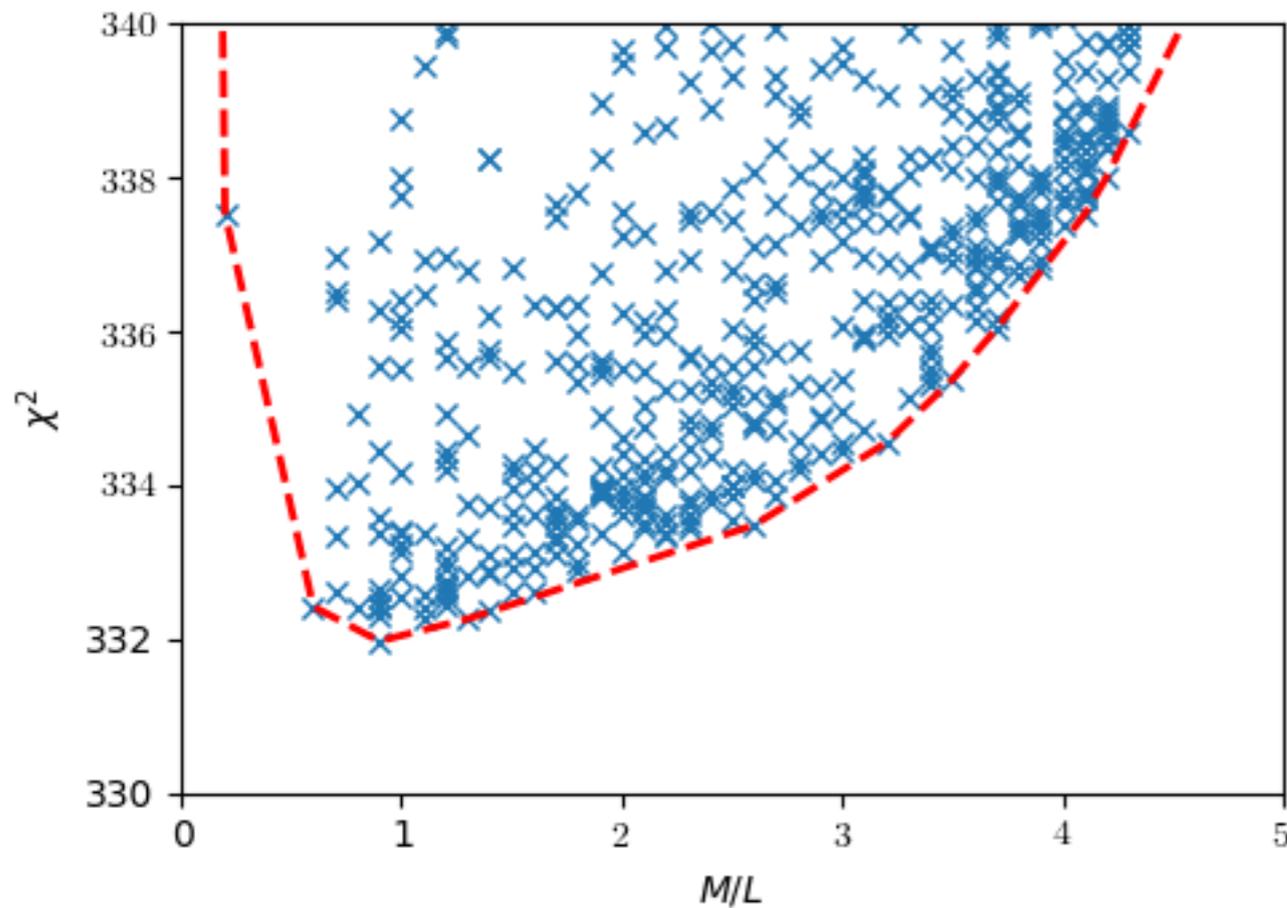
- Crowding is an issue in central regions
- We can't use individual velocity measurements

Updated velocity dispersion profile



- Central Mateo points suffer from crowding effects
- VIRUS-W measures a slight velocity rise towards the center

Results from Schwarzschild modeling



- Best fit black hole mass is $> 10^6 M_{\text{sun}}$
- Best fit M/L is lower than for previous models. The shape of the dark matter halo is unconstrained.

Conclusions

- Leo I shows a central density cusp. Steeper than predictions from DM cosmological simulations
- The central velocity dispersion shows a clear rise towards the center
- Results from Schwarzschild modeling point to a large central BH and less dark matter than previously measured